

# $p_T$ interpolation update

NPG Meeting

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# $p_T$ interpolation update

- Use results from previous experiments.
- Fit the  $p_T$  spectra and extract parameters A and T using parameterization  $\frac{Ap_T}{e^{p_T/T}+1}$ .
- From the three A and T parameters:
  - Interpolate to get A and T for  $\sqrt{s_{NN}} = 2.76$  TeV.
  - Interpolate also contours from error ellipse.

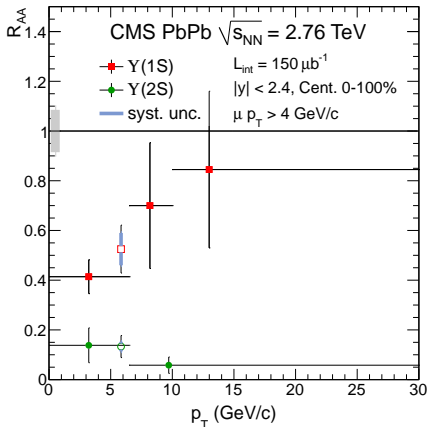
Main goals for this interpolation:

- Use for the  $R_{AA}$  vs  $p_T$  at  $\sqrt{s_{NN}} = 2.76$  TeV.
- Use as a reference for  $\sqrt{s_{NN}} = 5.00$  TeV.
- Use to cross check Ramona's theoretical calculations.
- Ramona's results status
  - Working with tuning the mass of the b quark.

# $R_{AA}$ vs $p_T$ using pp data for $\sqrt{s} = 2.76$ TeV

Example (Results have large uncertainties:)

$p_T \leq 30 \text{ GeV}/c$  and  $-2.4 \leq y \leq 2.4$



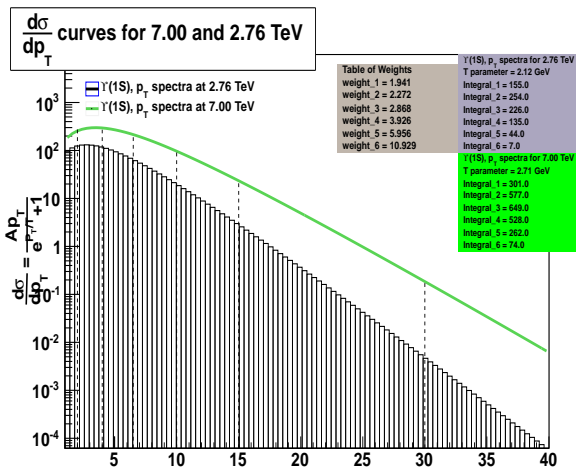
$ y $	$p_T$ [GeV/c]	$\langle p_T \rangle$ [GeV/c]	$\frac{d\sigma}{dy}$ [nb]
0.0-2.4	0-6.5	2.82	$0.668 \pm 0.091 \pm 0.115 \pm 0.040$
	6.5-10	8.36	$0.102 \pm 0.031 \pm 0.018 \pm 0.006$
	10-20	13.04	$0.037 \pm 0.013 \pm 0.006 \pm 0.002$
	0-20	4.73	$0.764 \pm 0.089 \pm 0.131 \pm 0.046$
0.0-1.2	0-20	5.18	$0.921 \pm 0.128 \pm 0.157 \pm 0.055$
1.2-2.4		4.03	$0.586 \pm 0.125 \pm 0.101 \pm 0.035$

**Table 13.** Cross section per unit of rapidity of Y(1S) as a function of rapidity and  $p_T$  in pp collisions. The average  $p_T$  value for each bin is given. Listed uncertainties are statistical first, systematic second, and global scale third. The latter is the uncertainty on the pp integrated luminosity.

# $p_T$ from A and T parameters for both energies

## Example (Weight definition:)

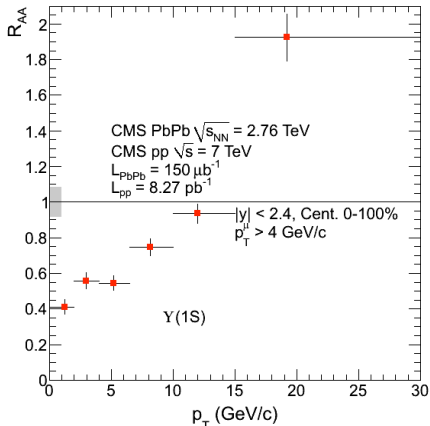
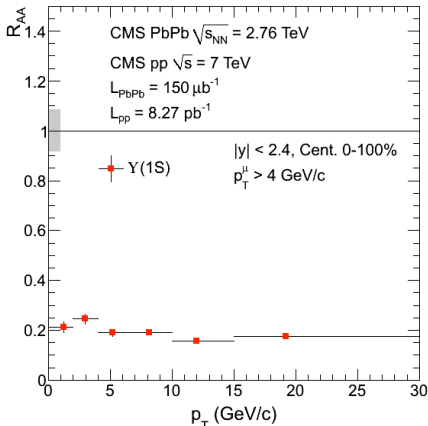
(Weight =  $\int_{lo}^{hi} (\sqrt{s_{pp}} = 7 \text{ TeV}) / \int_{lo}^{hi} (\sqrt{s_{pp}} = 2.76 \text{ TeV})$  where  $[lo, hi] = [0,2], [2,4], \dots$ )



# $R_{AA}$ vs $p_T$ before and after the Weights

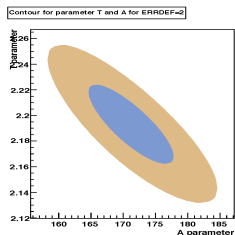
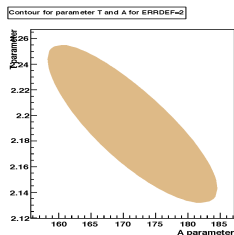
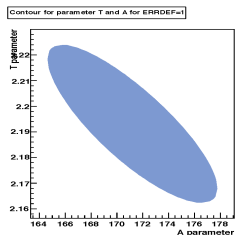
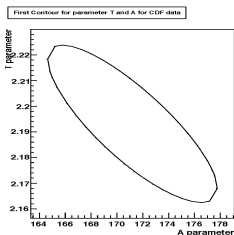
Example (Weight definition:)

(Weight =  $\int_{lo}^{hi} (\sqrt{s_{pp} = 7 \text{ TeV}}) / \int_{lo}^{hi} (\sqrt{s_{pp} = 2.76 \text{ TeV}})$  where  $[lo, hi] = [0, 2], [2, 4], \dots$ )

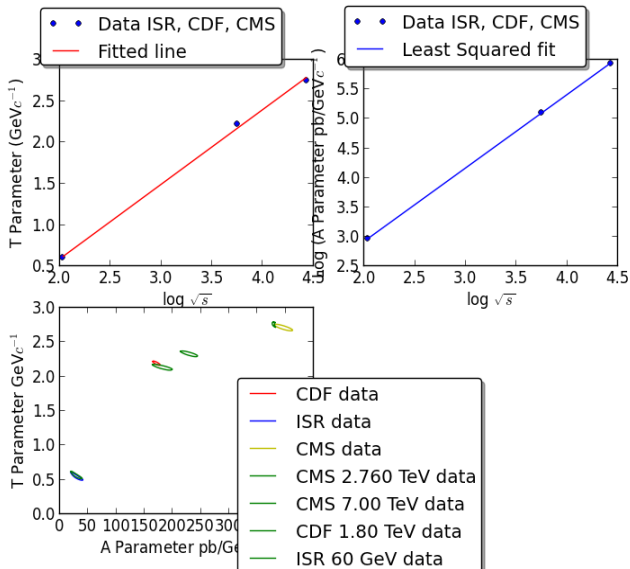


# Systematics on the weights

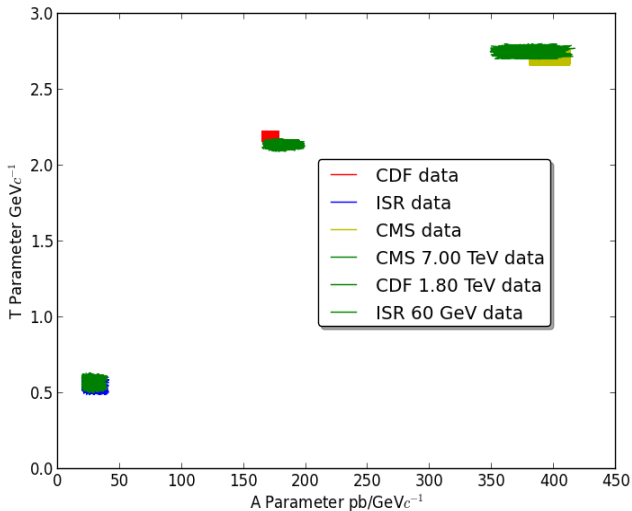
Start with one  $\sigma$  error ellipse for the three energies and perform same interpolation as for the center values. Go around the three ellipses and repeat process.



# Systematics on the weights: Interpolating ellipse



# Systematics on the weights: Randomizing 10000 points from each ellipse



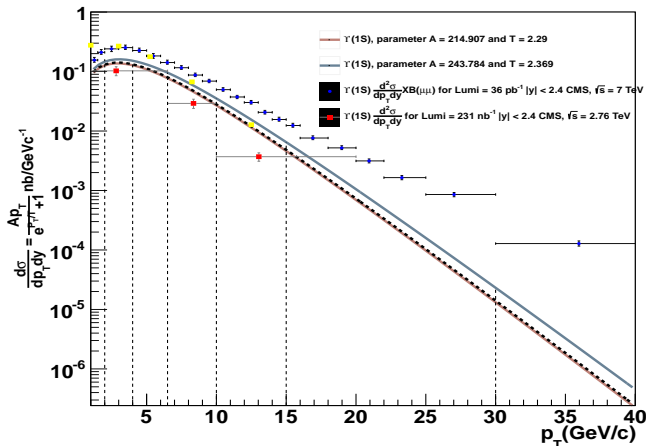


# $p_T$ spectra for $\sqrt{s} = 2.76$ TeV , $\sqrt{s} = 7.00$ TeV

Example (Notice branching ratio)

Extreme (anticorrelated) A and T parameters

## $p_T$ Spectra



- Interpolating using linear fits shifts error ellipse off center.
- In order to get proper coverage try different interpolation approach (non linear) or randomizing the  $\sigma$  contours.
- Cross check results with Ramona's calculation.
  - What should be the exact definition on the systematic error?
  - Is it just the ratio of the 7 TeV curve over the lowest/highest curve at 2.76 TeV or
  - Does the 7 TeV curve also have a lowest/highest curve.