

# Interference in vector meson production in Au+Au Collisions $\sqrt{s_{NN}} = 200$ GeV

- Review of analysis details
- Current status of results
- Latest alterations to the analysis
  - systematic checks
  - theory studies



Brooke Haag  
UC Davis

- Photonuclear Interaction
- Two nuclei “miss” each other ( $b > 2R_A$ ), electromagnetic interaction dominates over strong interaction
- Photon flux  $\sim Z^2$ 
  - Weizsäcker-Williams Equivalent Photon Approximation

$$\frac{d^3N(k,r)}{dkd^2r} = \frac{Z^2\alpha x^2}{\pi^2 kr^2} K_1^2(x)$$

$K_1(x)$  = Modified Bessel function

$k$  = photon energy

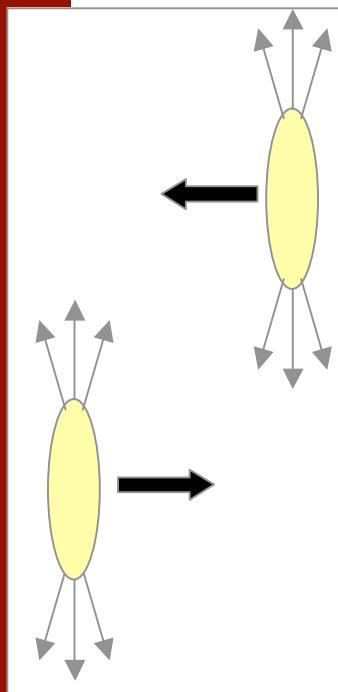
- No hadronic interactions

### $\rho^0$ Production

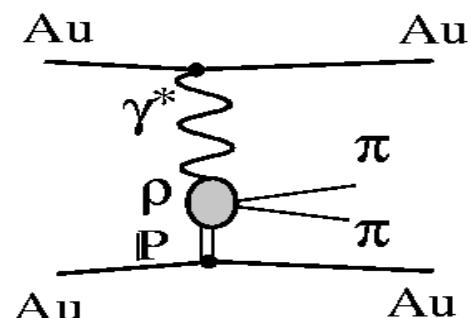
- Photon emitted by a nucleus fluctuates to virtual  $q\bar{q}$  pair
- Virtual  $q\bar{q}$  pair elastically scatters from other nucleus
- Real vector meson (i.e.  $J/\psi$ ,  $\rho^0$ ) emerges

### $\rho^0$ Production with coulomb excitation

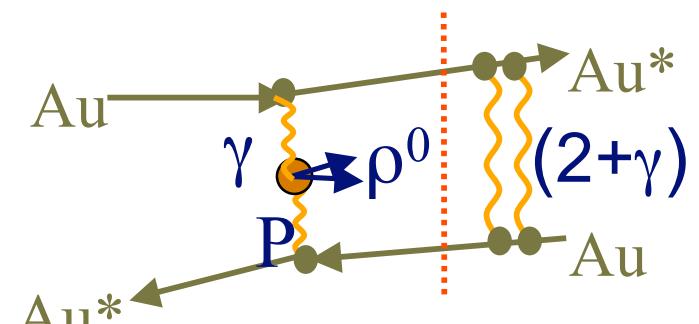
- Photons exchanged between ions give rise to excitation and subsequent neutron emission
- Process is independent of  $\rho^0$  production



## Ultra Peripheral Collisions



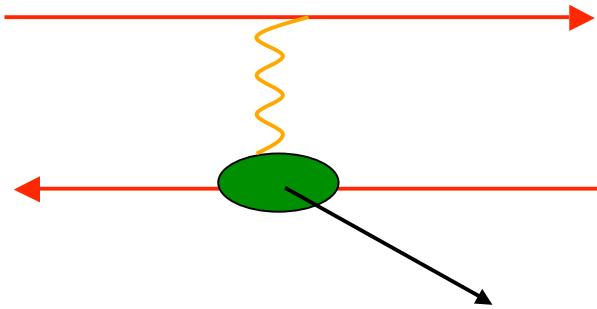
Courtesy of F. Meissner



Courtesy of S. Klein

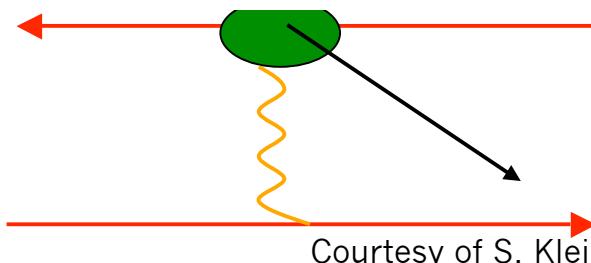
# Interference

Nucleus 1 emits photon which scatters from Nucleus 2



-Or-

Nucleus 2 emits photon which scatters from Nucleus 1



Courtesy of S. Klein

- Amplitude for observing vector meson at a distant point is the subtraction (since  $\rho$  parity is negative) of two plane waves:

$$A_o(x_o, \vec{p}, b) = A(p_\perp, y, b) e^{i[\phi(y) + \vec{p} \cdot (\vec{x} - \vec{x}_o)]} - A(p_\perp, -y, b) e^{i[\phi(-y) + \vec{p} \cdot (\vec{x} - \vec{x}_o)]}$$

- Cross section comes from square of amplitude:

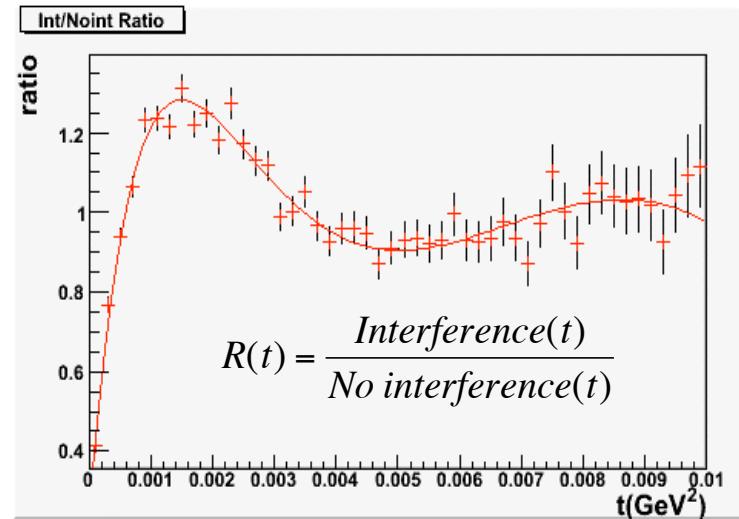
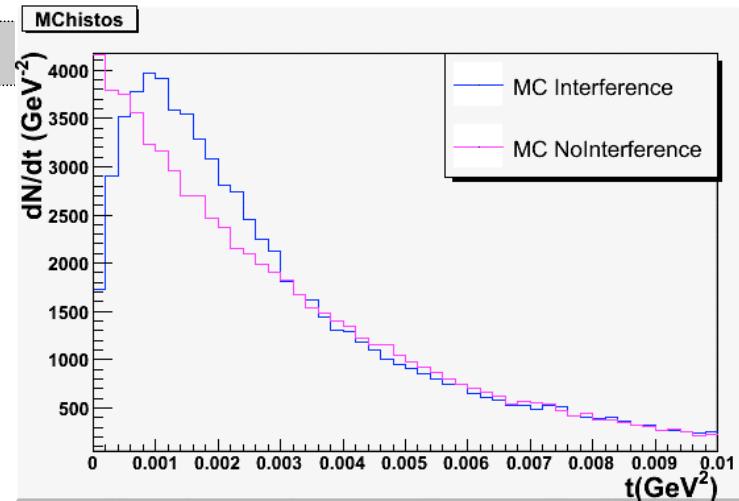
$$\sigma = A^2(p_\perp, y, b) + A^2(p_\perp, -y, b) - 2A(p_\perp, y, b)A(p_\perp, -y, b) \times \cos[\phi(y) - \phi(-y) + \vec{p} \cdot \vec{b}]$$

- We can simplify the expression if  $y \rightarrow 0$ :

$$\sigma = 2A^2(p_\perp, b)(1 - \cos[\vec{p} \cdot \vec{b}])$$

# Studying the Interference

- Generate MC  $t$  spectra with and without interference
- Calculate MC ratio in order to illustrate interference effect
- Fit MC ratio

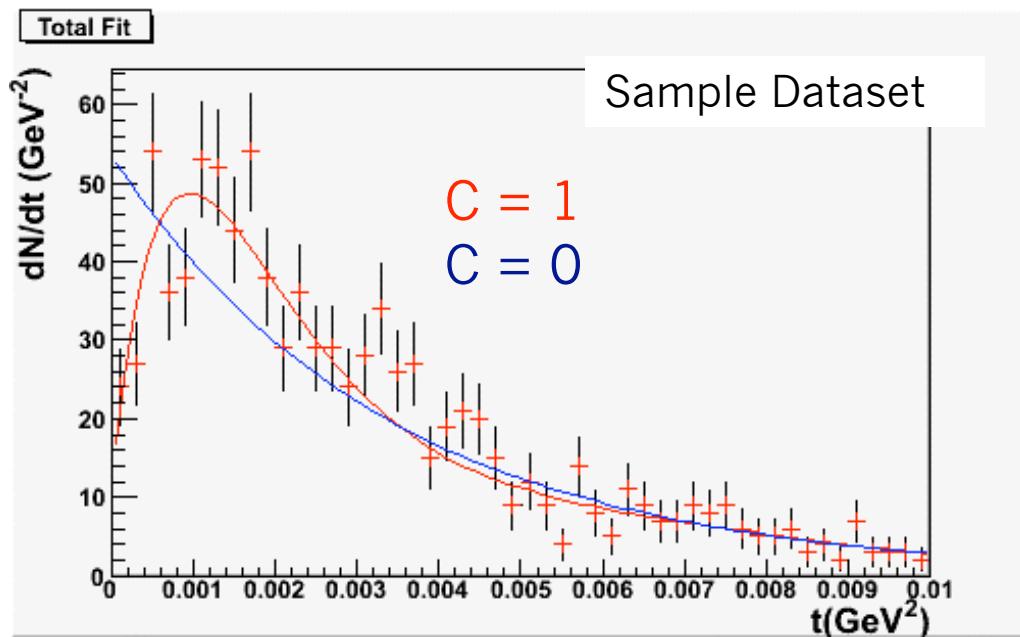


$$R(t) = a + \frac{b}{(t + 0.012)} + \frac{c}{(t + 0.012)^2} + \frac{d}{(t + 0.012)^3} + \frac{e}{(t + 0.012)^4}$$

# Measuring the Interference

- Apply overall fit

$$\frac{dN}{dt} = Ae^{-kt} (1 + c[R(t) - 1])$$



- A = overall normalization
- k = exponential slope
- c = degree of interference

$c = 1$

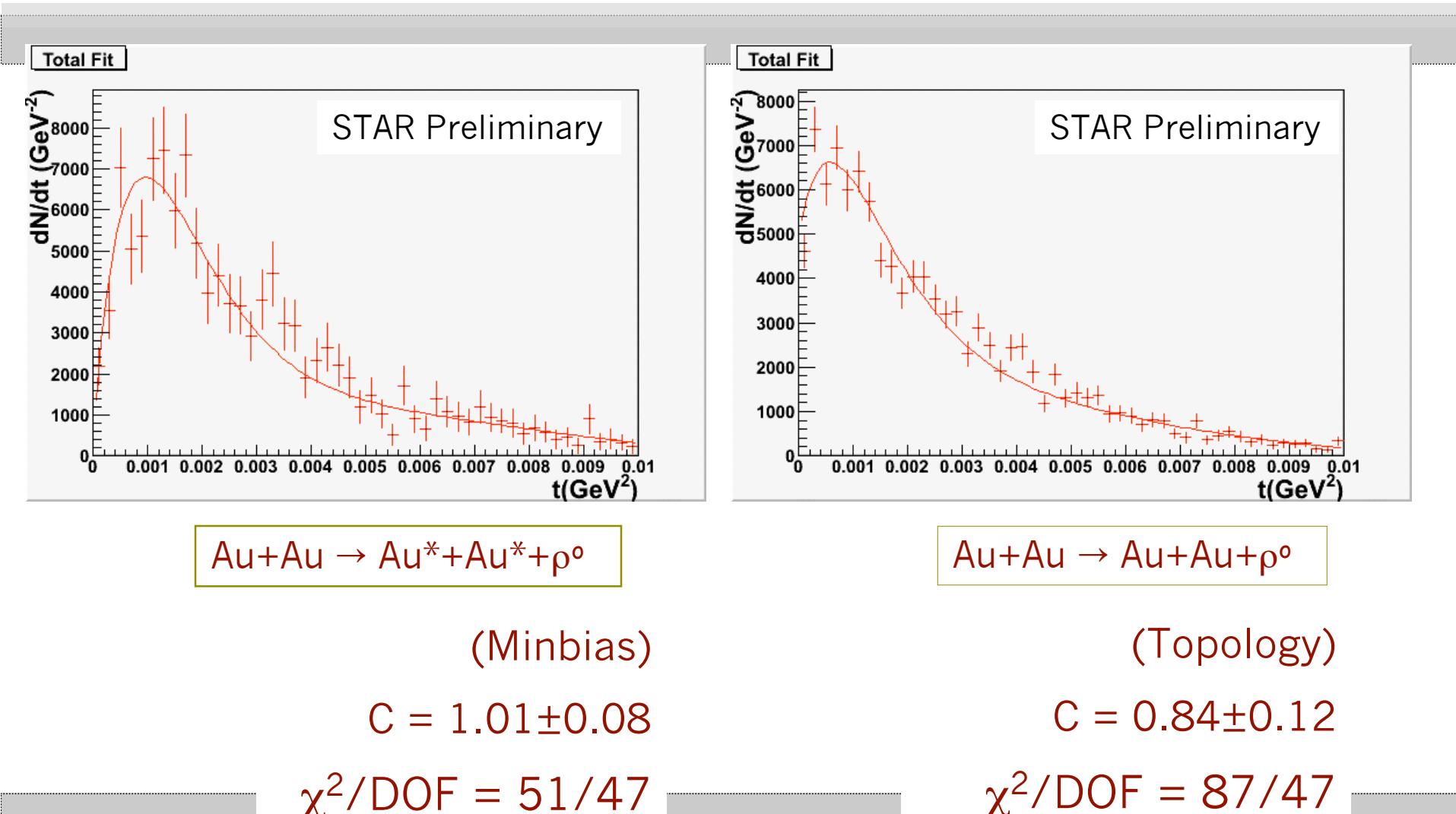
expected degree of interference

$c = 0$

no interference

Illustration of Fitting Methodology

# Results



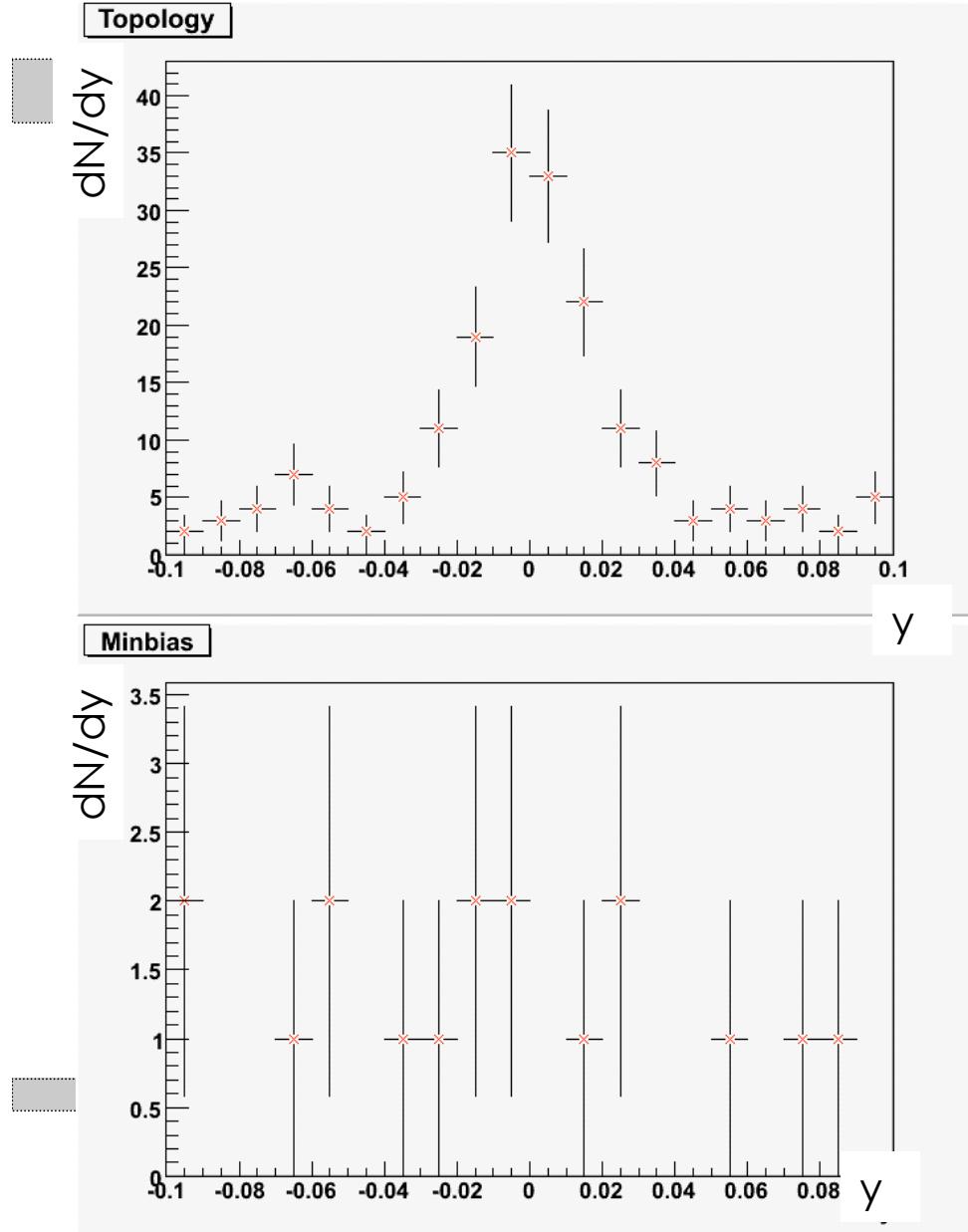
# Results Summary

	$C_{\text{preliminary}}$	$\chi^2/\text{dof}$ preliminary	$C$	$\chi^2/\text{dof}$
excitation				
$0 < y < 0.5$	$1.01 \pm 0.09$	50/47	$1.01 \pm 0.08$	51/47
$0.5 < y < 1.0$	$0.78 \pm 0.13$	72/47	$0.93 \pm 0.11$	80/47
No excitation				
$0.1 < y < 0.5$	$0.71 \pm 0.16$	81/47	$0.85 \pm 0.12$	88/47
$0.5 < y < 1.0$	$1.22 \pm 0.21$	50/47	$1.06 \pm 0.21$	84/47

# Latest Developments

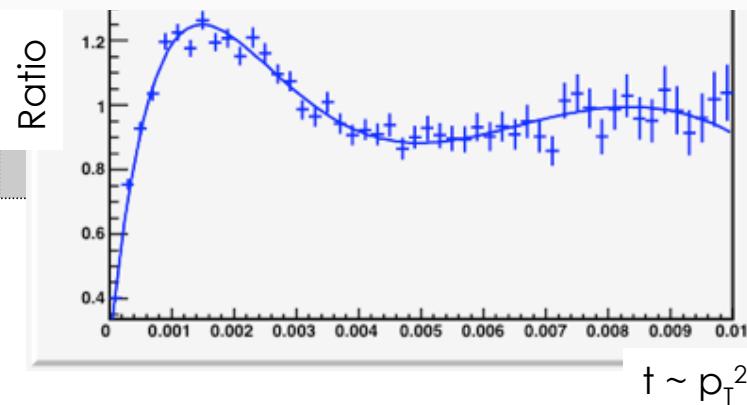
- Extended rapidity range for minbias analysis
- Systematic Error Studies
  - Fitting scheme
    - Better Fit for  $R(t)$
  - Theory comparisons
    - STARlight
    - KNLite - Adaptation of STARlight by Jim Draper

# Extended rapidity range

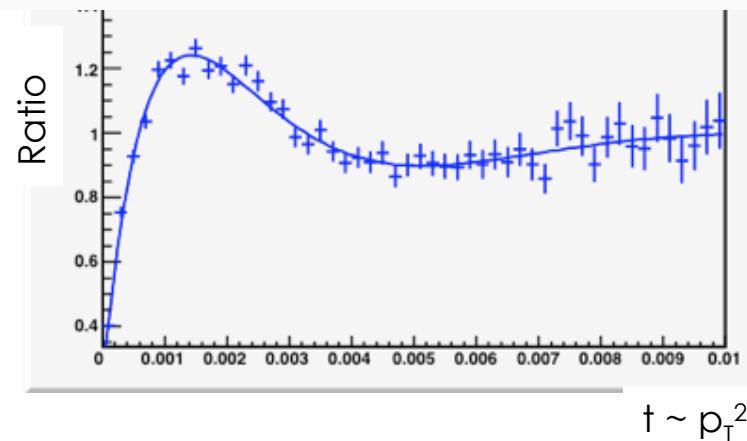


- Two rapidity ranges defined for the analysis
  - $0.1 < y < 0.5$
  - $0.5 < y < 1.0$
- Cut on midrapidity to eliminate cosmics
  - Effective for topology data, but unnecessary for minbias data

$$R(t) = a + \frac{b}{(t + 0.012)} + \frac{c}{(t + 0.012)^2} + \frac{d}{(t + 0.012)^3} + \frac{e}{(t + 0.012)^4}$$

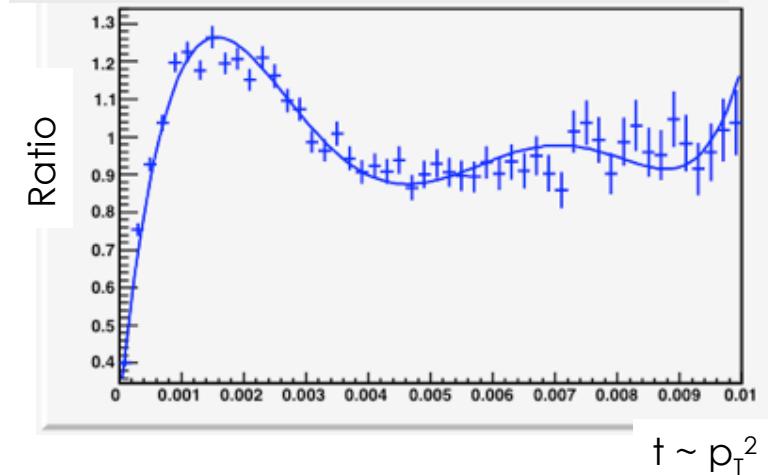


$$R(t) = a + \frac{b}{(t + 0.012)} + \frac{c}{(t + 0.012)^2} + \frac{d}{(t + 0.012)^3} + \frac{e}{(t + 0.012)^4} + \frac{f}{(t + 0.012)^5}$$

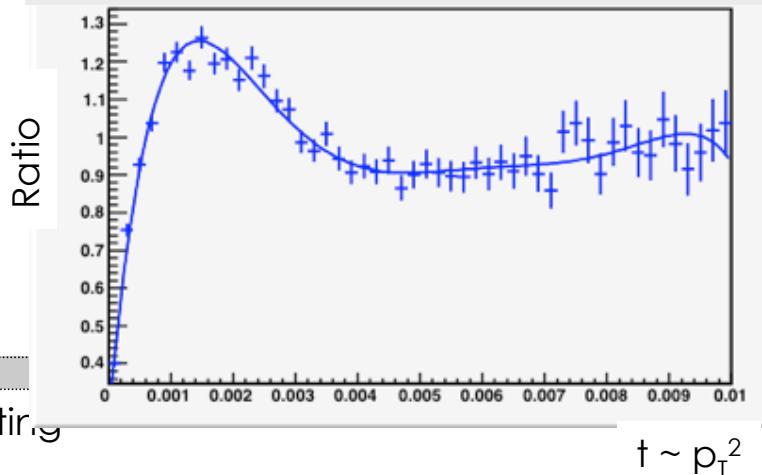


## Fitting $R(t)$

$$R(t) = at^5 + bt^4 + ct^3 + dt^2 + et + f$$



$$R(t) = at^6 + bt^5 + ct^4 + dt^3 + et^2 + ft + g$$



$$\frac{dN}{dt} = Ae^{-kt}(1 + c[R(t) - 1])$$

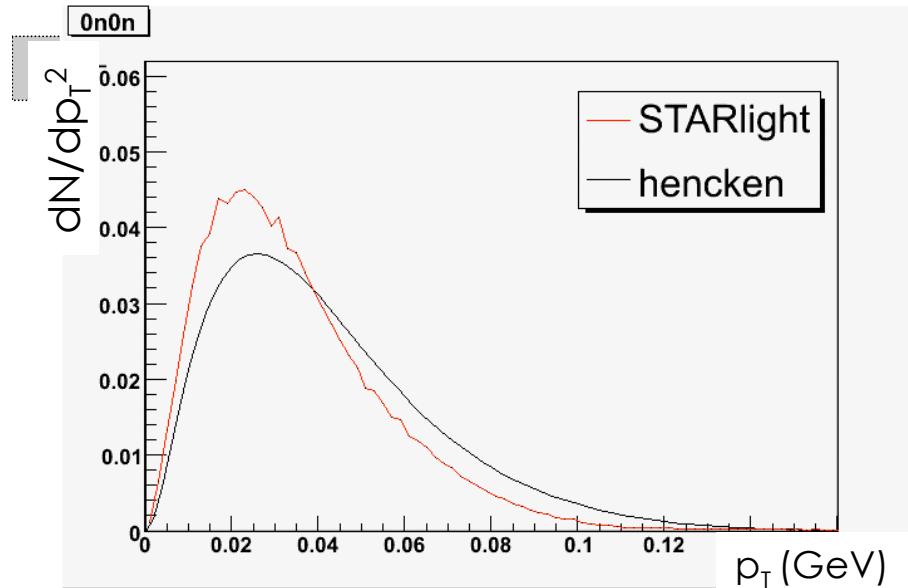
The overall fit is:

We develop an estimate of the systematic error due to the fit by trying different fit functions.

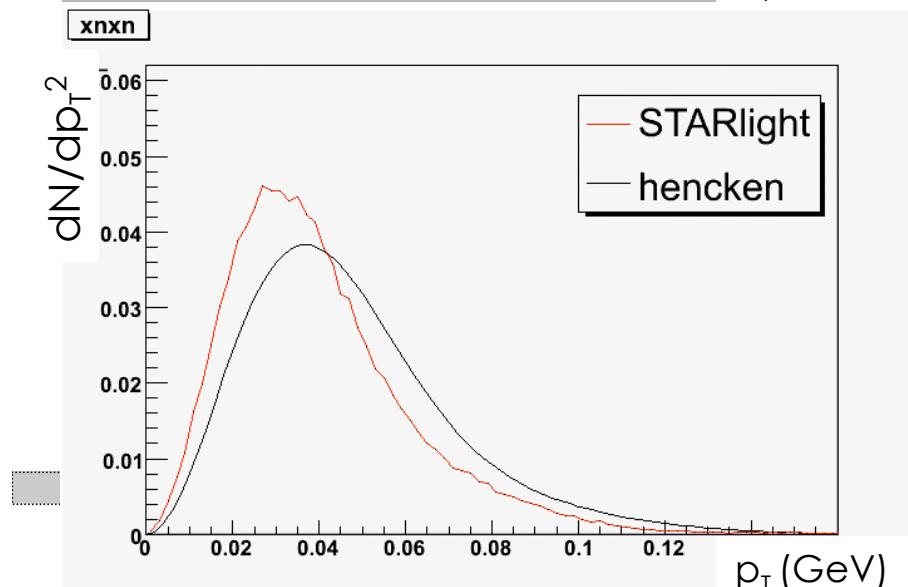
# Theory Comparisons

- 2 main theories describing interference:
  - STARlight - S. Klein and J. Nystrand
  - Hencken, Baur, Trautmann (HBT)  
PRL96(2006)012303
- New model KNLite, adaptation of STARlight - Jim Draper
  - Better understanding of  $R(t)$  [interference] out to 300 MeV
  - Better match to HBT
  - Studies of sensitivity to Nuclear Radius
  - Studies of sensitivity to  $m_p$

# Theory Comparisons - STARlight

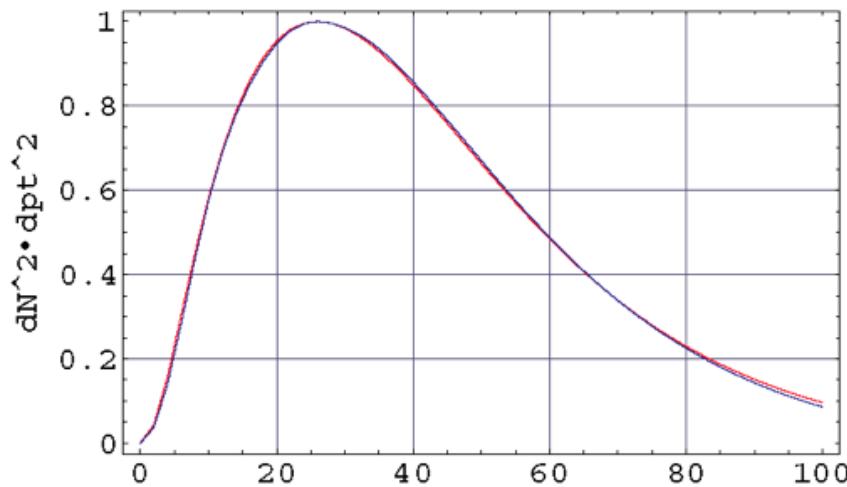


- STARlight vs. HBT predictions
  - Top = no nuclear excitation
  - Bottom = multiple nuclear excitation

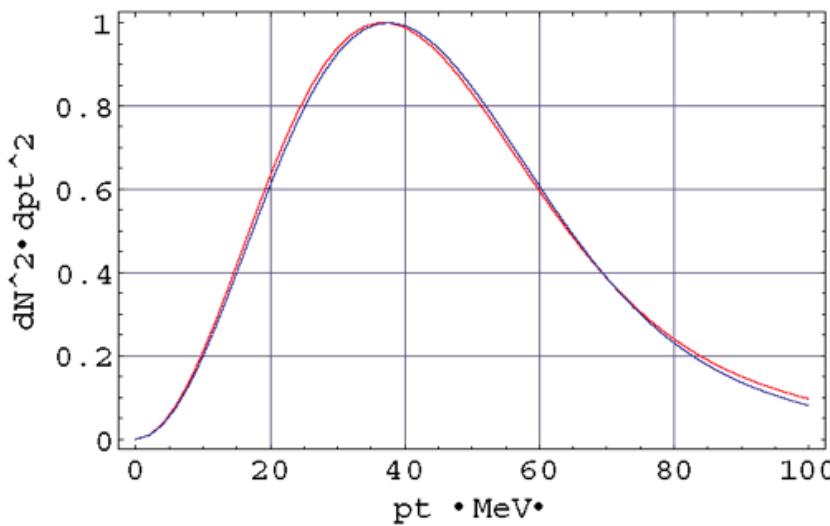


# Theory Comparisons - KNLite

R6.7 Intrfrnc 0tag HBT• red• KNLite• blue•

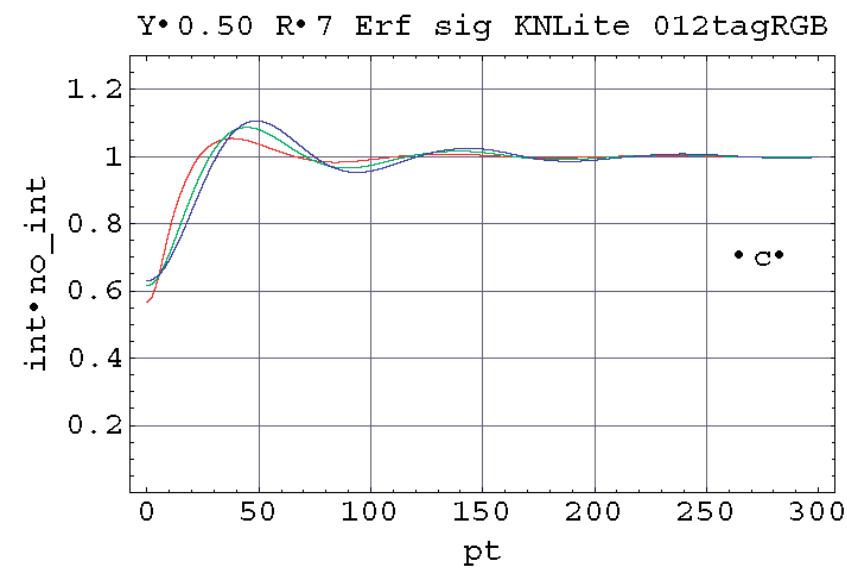
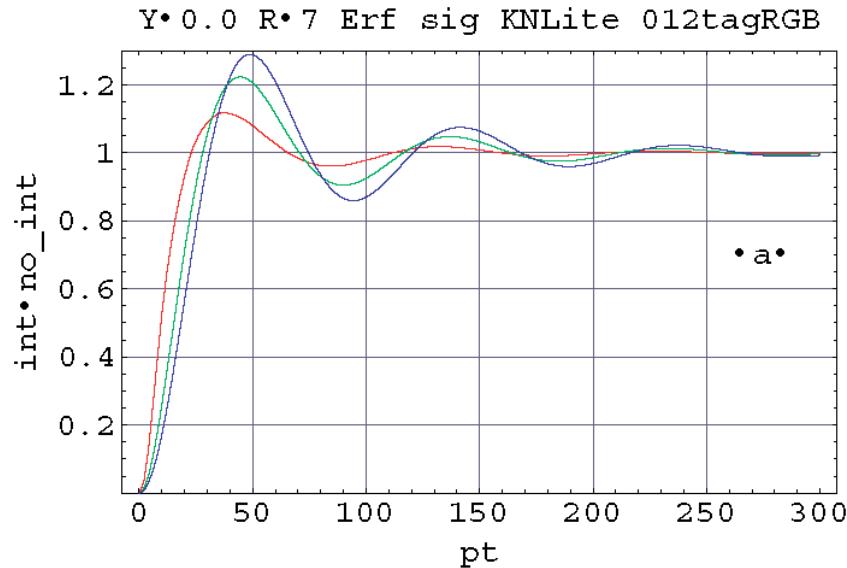


R6.7 Intrfrnc 2tag HBT• red• KNLite• blue•



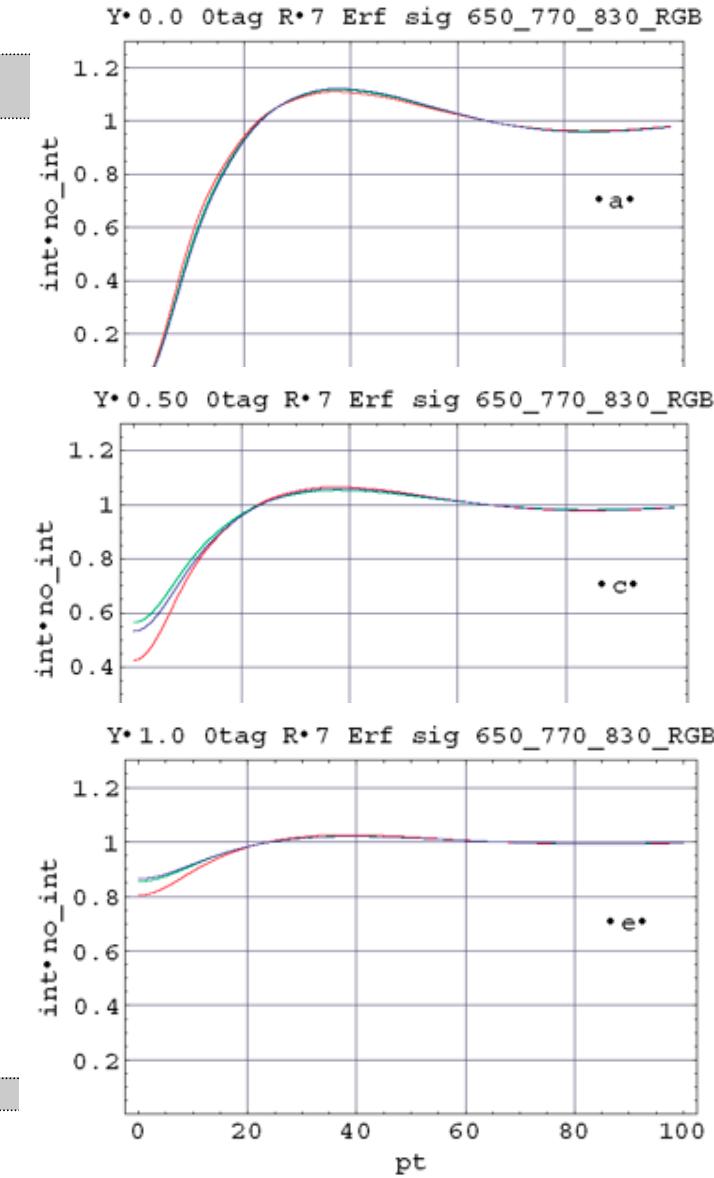
- KNLite vs HBT predictions
  - HBT = red, KNLite = blue
  - Top = no nuclear excitation
  - Bottom = multiple nuclear excitation

# Theory Comparisons - KNLite



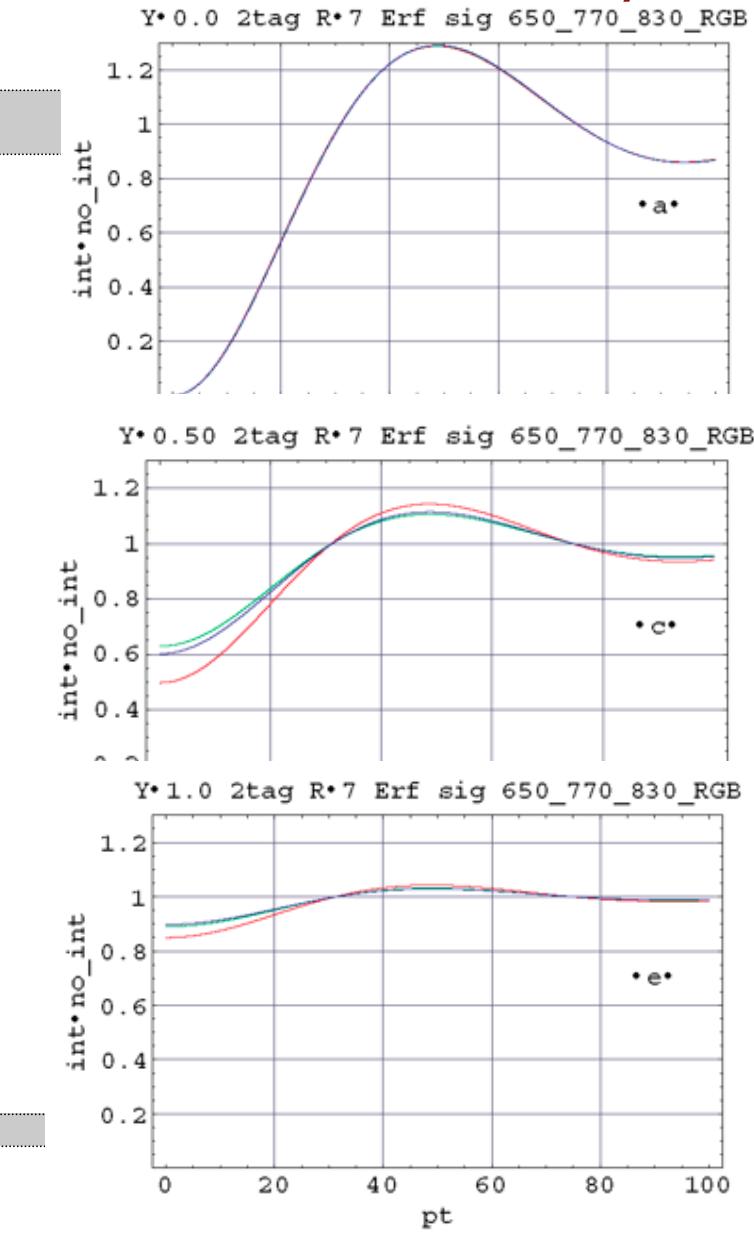
- Ratio of Interference to No Interference  $p_T$  spectra
  - Two rapidity bins:  $y = 0$ ,  $y = 0.5$
  - Red = no nuclear excitation
  - Green = single excitation
  - Blue = double excitation
- Clear evidence of interference out to 300 MeV

# Theory Comparisons - KNLite



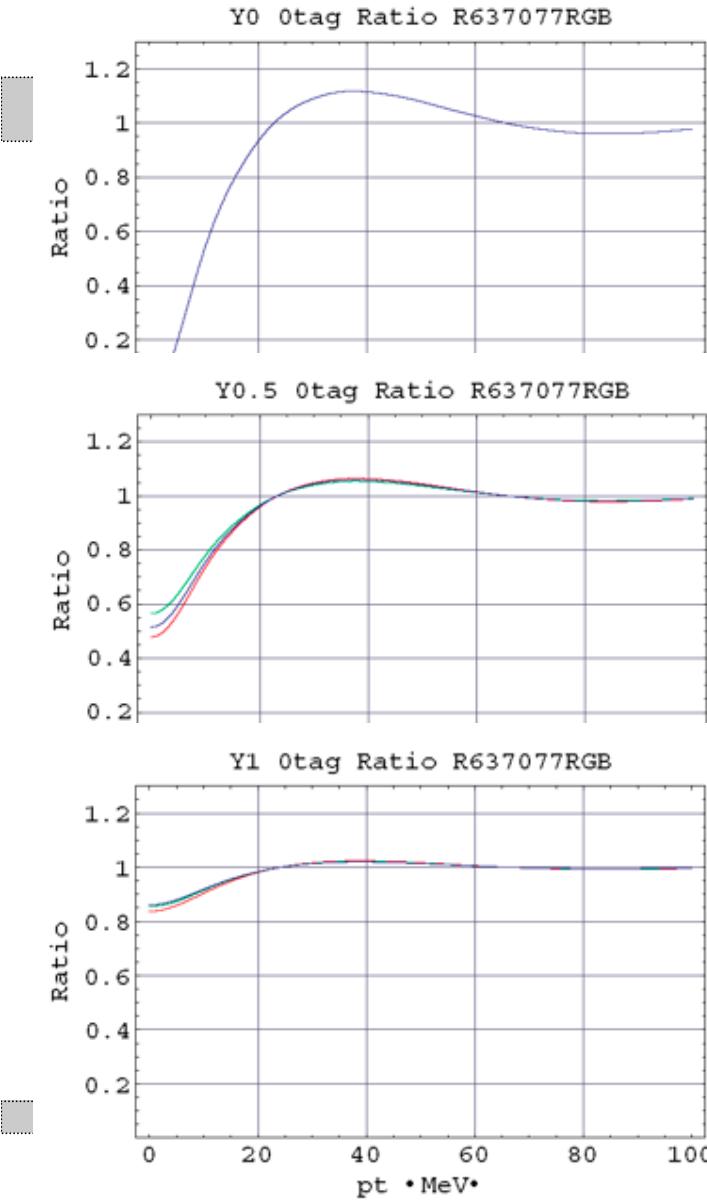
- Ratio of Interference to No Interference  $p_T$  spectra - mass dependence
  - no nuclear excitation
  - Top -  $y = 0$
  - Middle -  $y = 0.5$
  - Bottom -  $y = 1.0$
  - Three mass assumptions
    - Red =  $\rho^0$  mass = 650 MeV
    - Green =  $\rho^0$  mass = 770 MeV
    - Blue =  $\rho^0$  mass = 830 MeV

# Theory Comparisons - KNLite



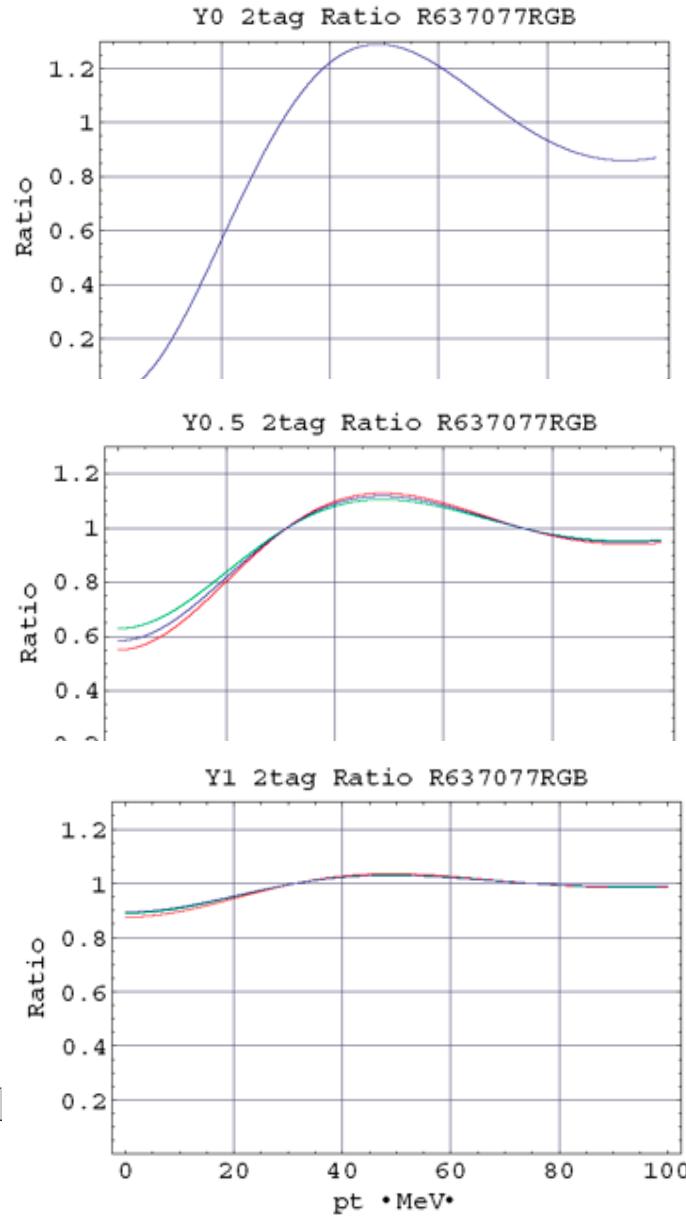
- Ratio of Interference to No Interference  $p_T$  spectra - mass dependence
  - double nuclear excitation
  - Top -  $y = 0$
  - Middle -  $y = 0.5$
  - Bottom -  $y = 1.0$
  - Three mass assumptions
    - Red =  $\rho^0$  mass = 650 MeV
    - Green =  $\rho^0$  mass = 770 MeV
    - Blue =  $\rho^0$  mass = 830 MeV

# Theory Comparisons - KNLite



- Ratio of Interference to No Interference  $p_T$  spectra - nuclear radius dependence
  - no nuclear excitation
  - Top -  $y = 0$
  - Middle -  $y = 0.5$
  - Bottom -  $y = 1.0$
  - Three radius assumptions
    - Red = 6.3 fm
    - Green = 7.0 fm
    - Blue = 7.7 fm

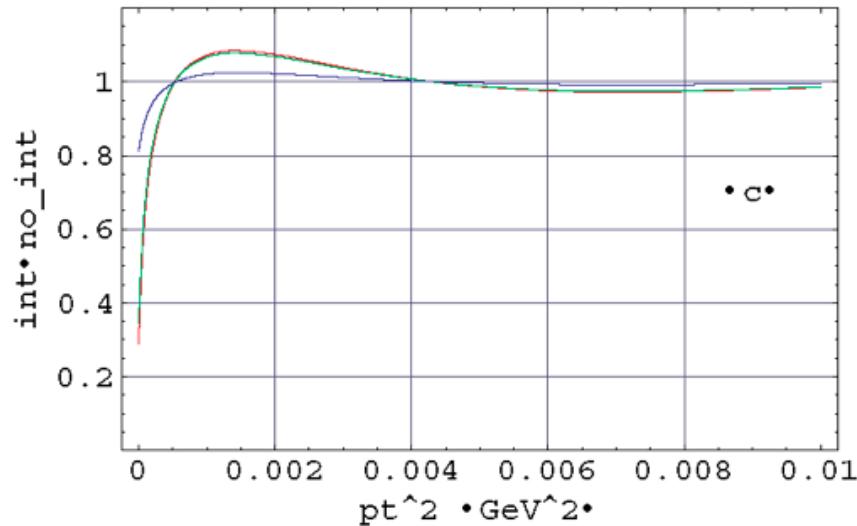
# Theory Comparisons - KNLite



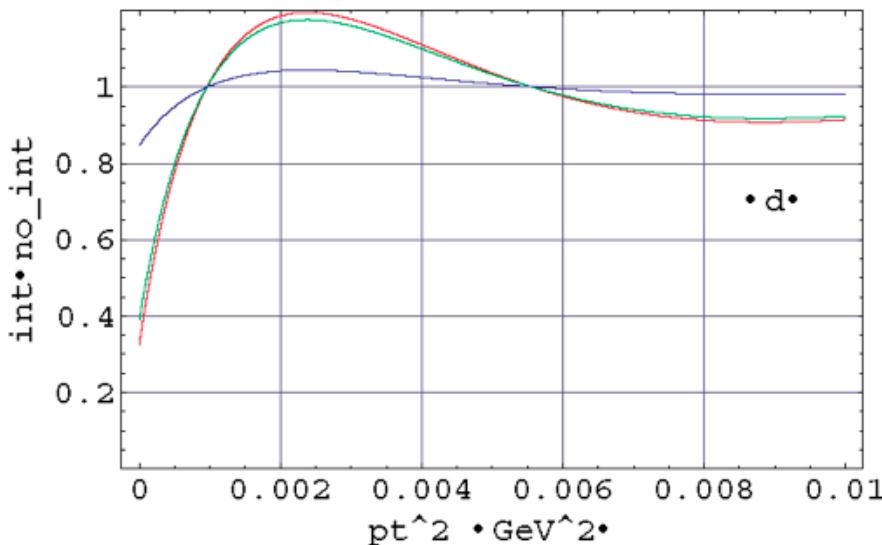
- Ratio of Interference to No Interference  $p_T$  spectra - nuclear radius dependence
  - double nuclear excitation
  - Top -  $y = 0$
  - Middle -  $y = 0.5$
  - Bottom -  $y = 1.0$
  - Three radius assumptions
    - Red = 6.3 fm
    - Green = 7.0 fm
    - Blue = 7.7 fm

# Theory Comparisons - KNLite

0tag: Y0.0to0.5 Y0.1to0.5 Y0.5to1.0 RGB



2tag: Y0.0to0.5 Y0.1to0.5 Y0.5to1.0 RGB



- Ratio of Interference to No Interference  $p_T^2$  spectra
  - Top = no nuclear excitation
  - Bottom = double excitation
  - Three rapidity ranges
    - Red =  $0.0 < y < 0.5$
    - Green =  $0.1 < y < 0.5$
    - Blue =  $0.5 < y < 1.0$

# Conclusions and Outlook

- Fitting scheme well refined
  - Still trouble with  $\chi^2$  on several of the fits
    - considering scaling data by  $\sqrt{\chi^2}$  as outlined in particle data guide
- Good handle on systematics
  - Theory has provided biggest trouble
  - KNLite provides better idea of systematics and better comparisons to established theory
  - Likely KNLite will be used to define  $R(t)$  in the final data fits
- Paper draft currently being worked on