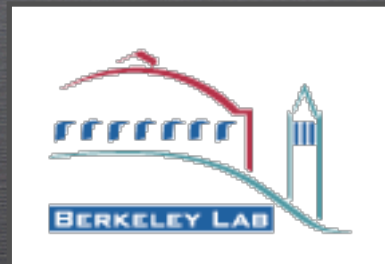


# GHOSTLY HEAT

## UNDERSTANDING NEUTRINOS THROUGH CALORIMETRY

THOMAS D. GUTIERREZ  
LAWRENCE BERKELEY NATIONAL LABORATORY



CAL POLY PHYSICS COLLOQUIUM, FEB 2006

CAL POLY



# OVERVIEW

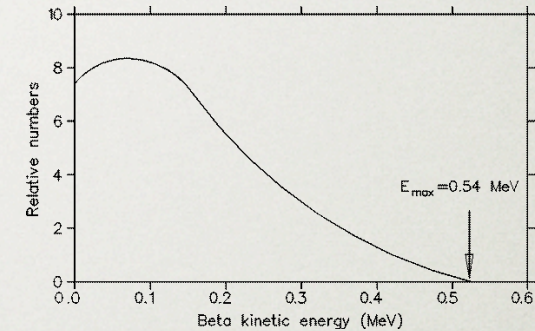
---

- What are neutrinos?
- How do we know neutrinos have mass? What is that mass?
- Are neutrinos their own antiparticle?
- Why are these questions important?
- How can measuring an exotic nuclear decay help us?
- How do you use calorimetry to measure this decay?
- What experiment uses this method?
- How students can get involved and at what level?



# BETA DECAY

Beta decay seems to violate energy-momentum conservation



1930

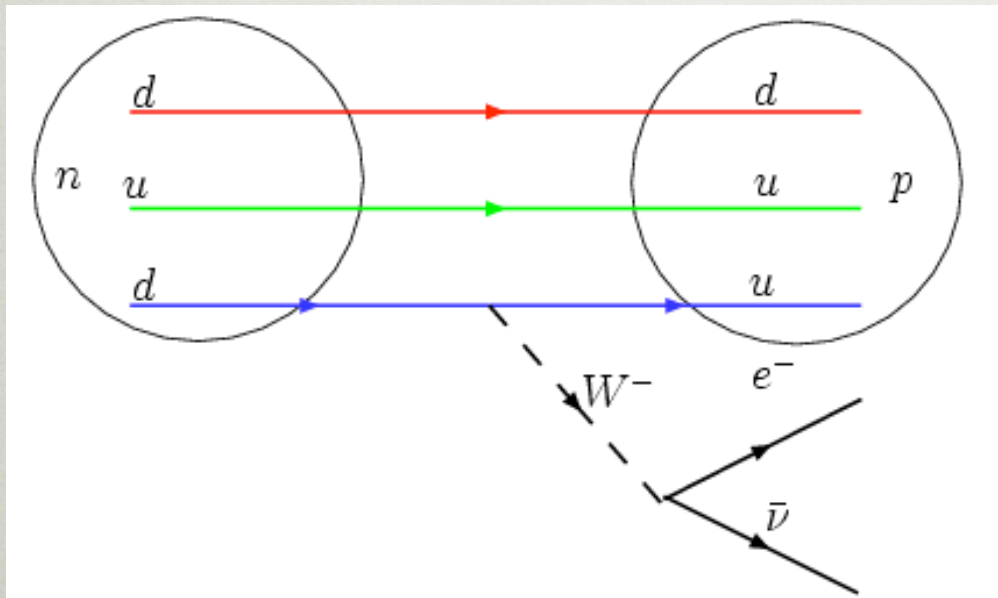
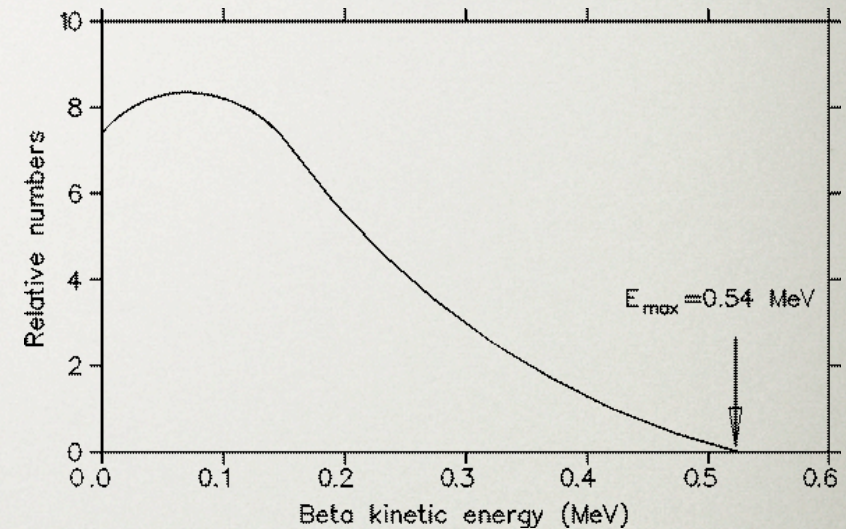


Liebe Radioaktive  
Damen und Herren



# THE NEW BETA DECAY

A neutron is converted to a proton via a weak force interaction



A three body final state that includes a massless fermion naturally explains the continuum of electron energies

1933



Lo chiamo il neutrino



# NEUTRINOS INTERACT WEAKLY

They are often called “ghost particles” because they only interact weakly; they are notoriously difficult to measure

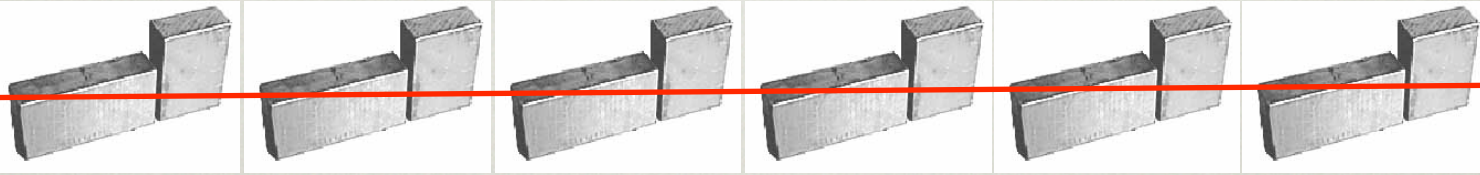
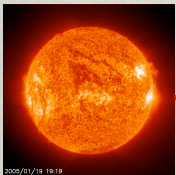
The sun produces  
copious electron  
neutrinos

$\sim 10^{11}$  solar neutrinos per  
second pass through your  
thumb

Despite being difficult to  
measure, it was eventually  
discovered in 1956

1 light year of solid lead

1/2 neutrinos  
from the sun still  
get through

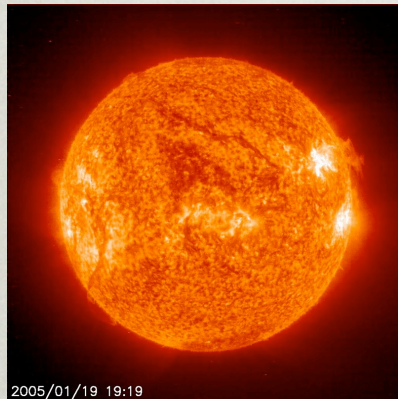




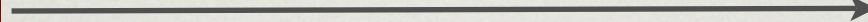
# THE SOLAR NEUTRINO PROBLEM

Careful measurements of electron neutrinos from the sun showed a substantial deficit compared to otherwise established solar models

Only  $\sim 1/3$  of the electron neutrinos predicted by Bahcall's model were measured by Davis

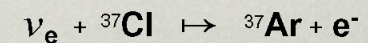


$\nu_e$



$\nu_e$

1964: Davis and Bahcall  
at Homestake





# THE STANDARD MODEL

In the Standard Model, the neutrino is a massless neutral fermion with several important properties

FERMIONS			matter constituents spin = 1/2, 3/2, 5/2, ...		
Leptons spin = 1/2			Quarks spin = 1/2		
Flavor	Mass GeV/c <sup>2</sup>	Electric charge	Flavor	Approx. Mass GeV/c <sup>2</sup>	Electric charge
$\nu_e$ electron neutrino	$<1 \times 10^{-8}$	0	<b>u</b> up	0.003	2/3
<b>e</b> electron	0.000511	-1	<b>d</b> down	0.006	-1/3
$\nu_\mu$ muon neutrino	$<0.0002$	0	<b>C</b> charm	1.3	2/3
<b><math>\mu</math></b> muon	0.106	-1	<b>S</b> strange	0.1	-1/3
$\nu_\tau$ tau neutrino	$<0.02$	0	<b>t</b> top	175	2/3
<b><math>\tau</math></b> tau	1.7771	-1	<b>b</b> bottom	4.3	-1/3

Comes in 3 different flavors

The flavors (“electron”, “muon”, “tau”) are regarded as fundamental quantum numbers in their own right

+Antiparticles

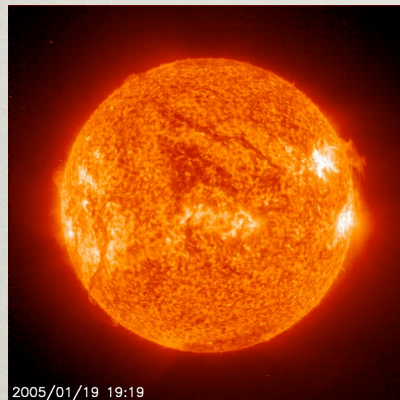


# PROBLEM SOLVED

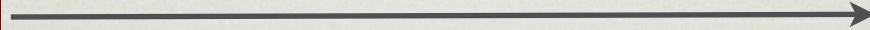
When all flavors are accounted for (~35 years later!), the number of total neutrinos from the sun matches Bahcall's model



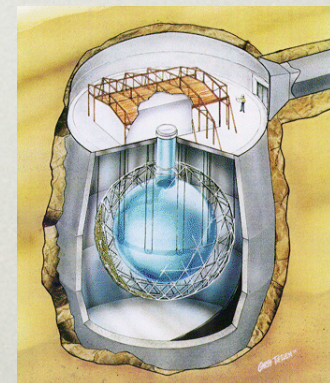
The SNO experiment accounted for all neutrino flavors; electron neutrinos must be transforming into other kinds of neutrinos mid-flight



$$\nu_e \quad \nu_\mu + \nu_e + \nu_\tau$$



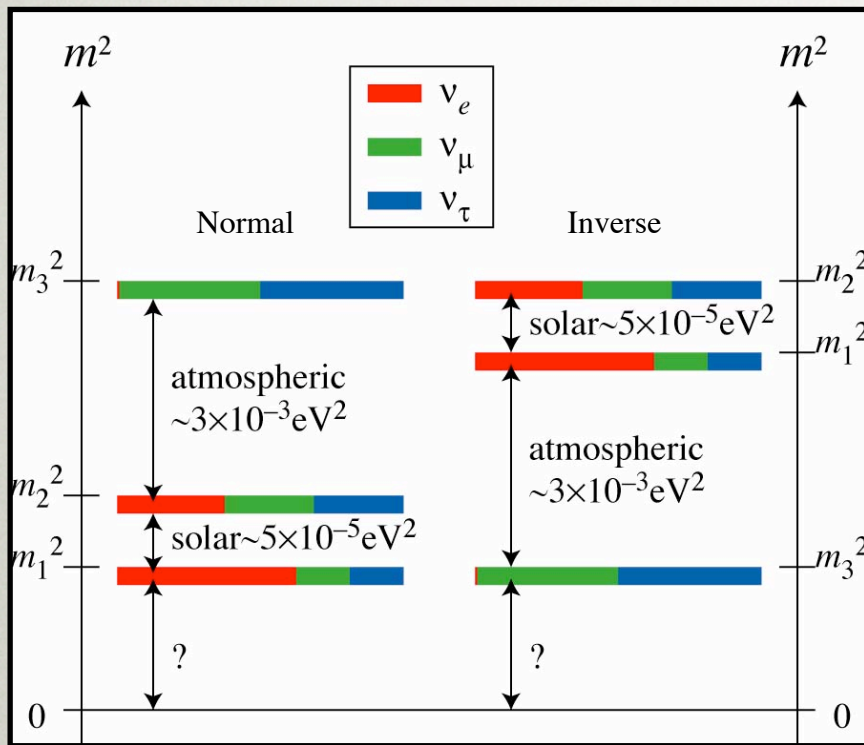
How can this be?





# FLAVOR OSCILLATIONS

The neutrinos we measure in reactions are not energy eigenstates



Neutrinos of definite mass (energy)  
don't change in time

We measure neutrinos of definite  
FLAVOR in the lab

Flavors are made of different mass  
states so do change in time

The absolute mass scale and ordering is  
unknown

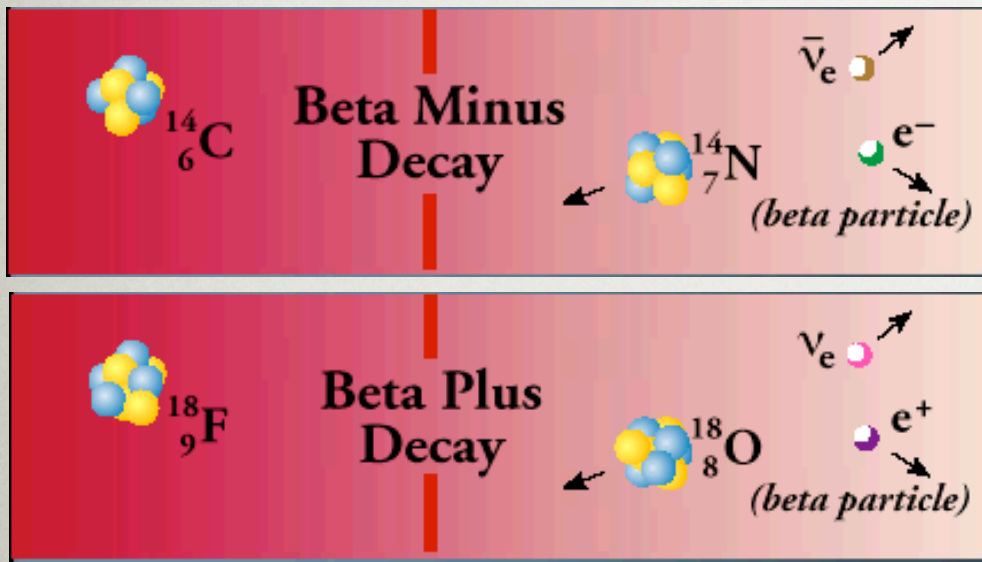


Neutrinos have mass but we don't know what it is



# THE ANTINEUTRINO

Can we really tell neutrinos apart from their antiparticles?



In contrast to other fermions, the neutrino has no electric charge or internal structure, so the distinction between a “particle” and “antiparticle” is subtle

The classification is by the type of reactions observed and then by following a convention

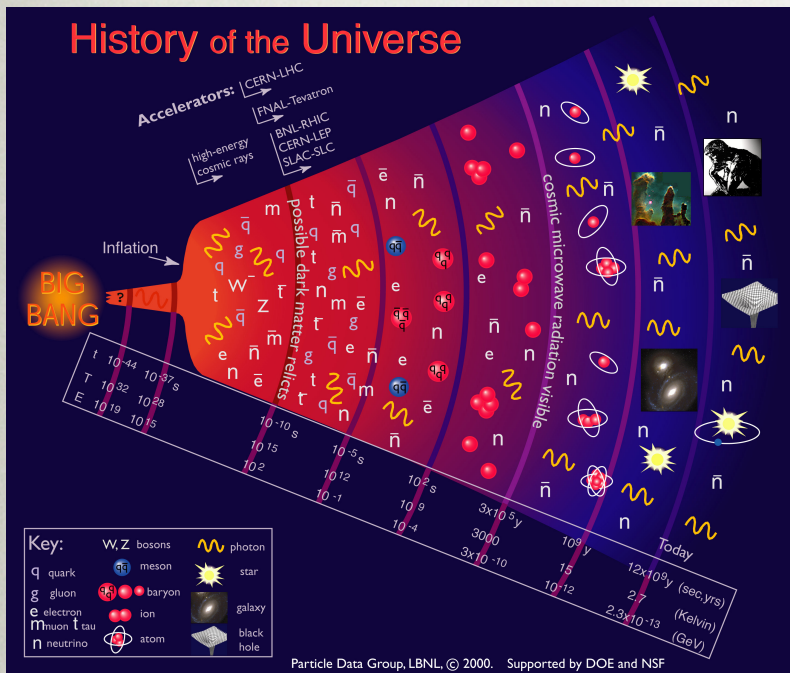


The neutrino could be it's own antiparticle



# WHY IS IT IMPORTANT?

Knowing the neutrino's mass and if it is distinct from its antiparticle may have profound implications for cosmology



Knowing the neutrino mass will provide the neutrino contribution to dark matter

Knowing the nature will help understand the matter/antimatter asymmetry in the universe

May provide footholds into other Beyond Standard Model physics



# Review and Onward



Experiments have shown:

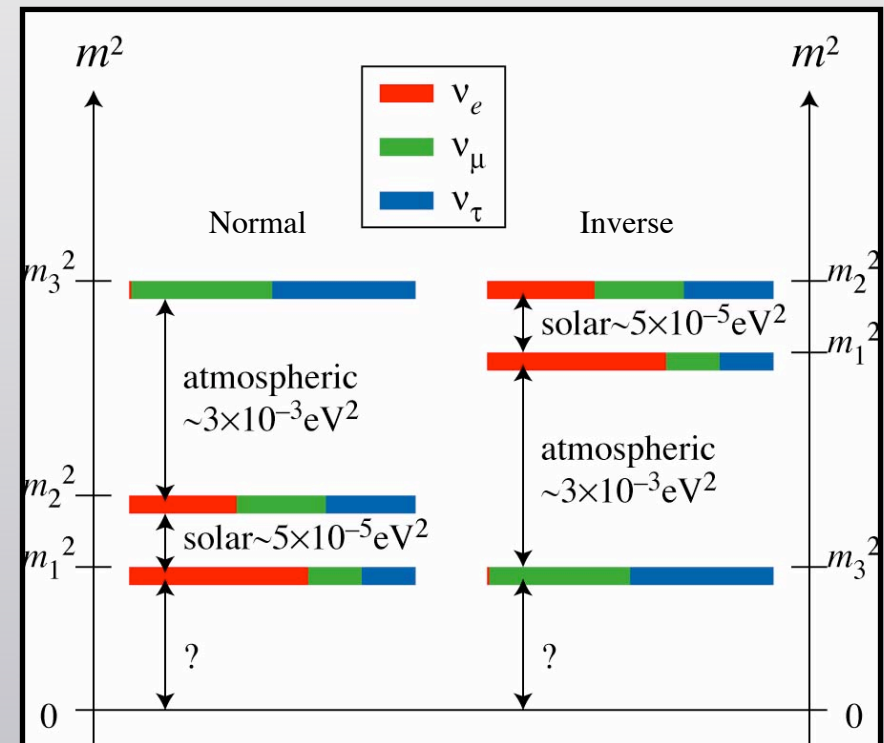
- Neutrinos undergo flavor-changing oscillations
  - This implies neutrinos have finite masses
  - Only sensitive to mass difference squared



Important open questions in  $\nu$  physics

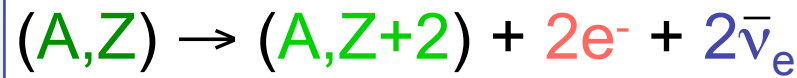
- What is absolute mass?
- What is correct mass ordering?
- Are they their own antiparticles?

Amazingly, there is a single measurement that can address these





# $\beta\beta 2\nu$ : Double Beta Decay



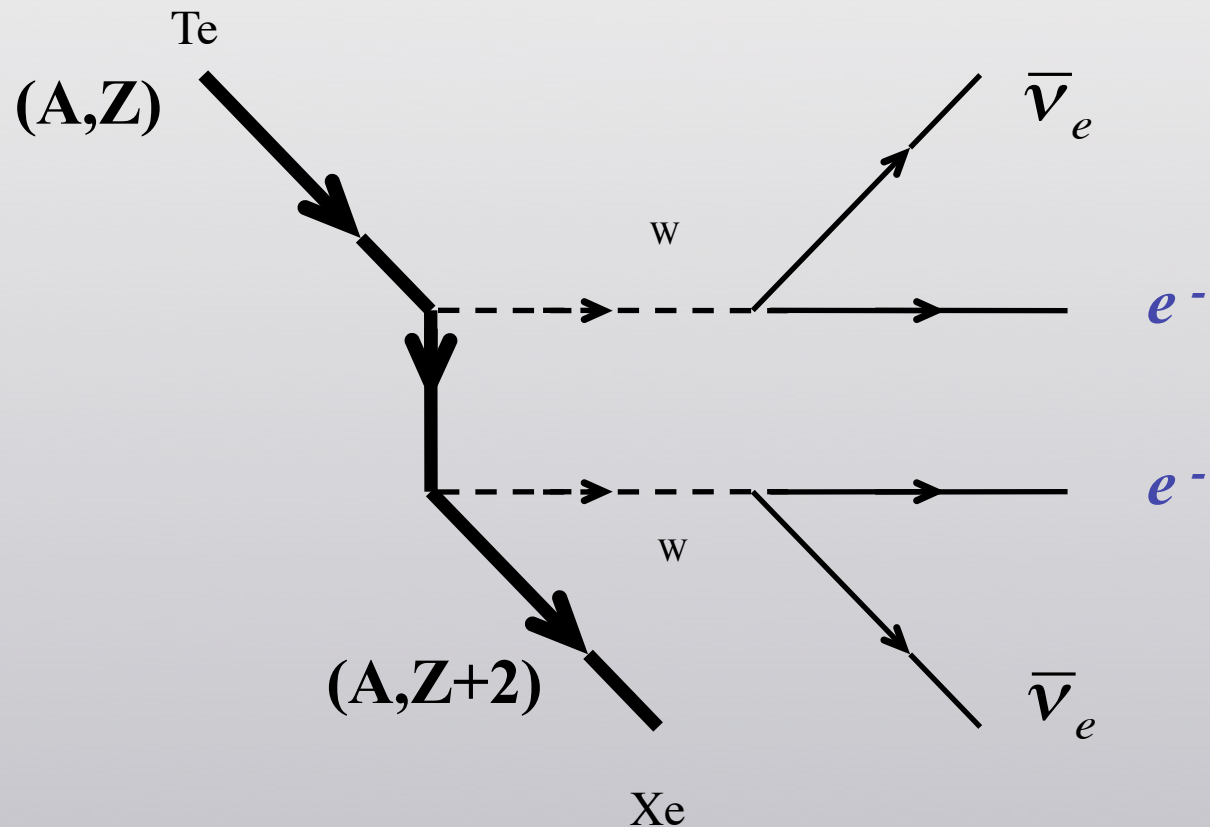
Really is just two simultaneous beta decays: very unlikely

Allowed by the Standard Model

$$\tau \geq 10^{18} \text{ y}$$

Measured in real systems  
(NEMO, geochemical, etc.)

Time scale reminder:  
NASA's WMAP mission  
(2003) sets the age of the  
universe at  
 $13.7 \pm 0.2 \times 10^9$  years





# $\beta\beta 0\nu$ : Neutrinoless Double Beta Decay

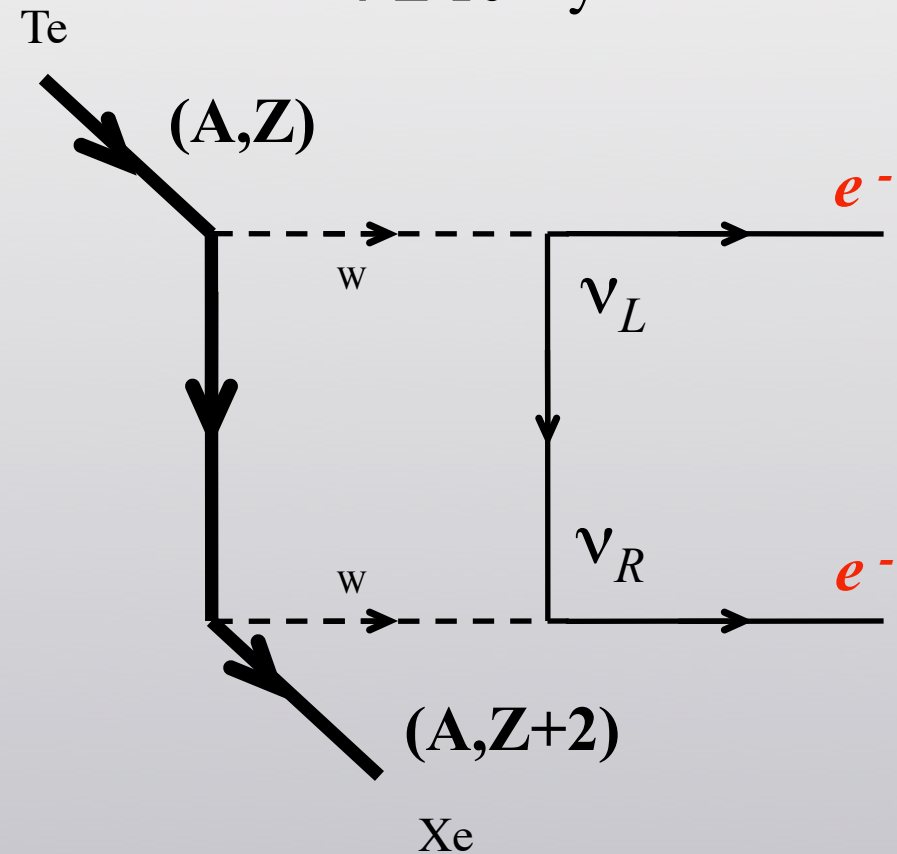


open discussion on its  
observation  
 $\tau \geq 10^{25}$  y

**Observation of  $\beta\beta 0\nu$  implies Physics  
beyond the Standard Model**

- Rate of decay sets  $\nu$  mass scale and reveals hierarchy
- Process only occurs if neutrinos are their own anti-particles

A long lifetime means you need more moles (thus mass) of source material to measure the decay in a reasonable time

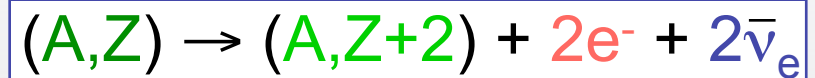
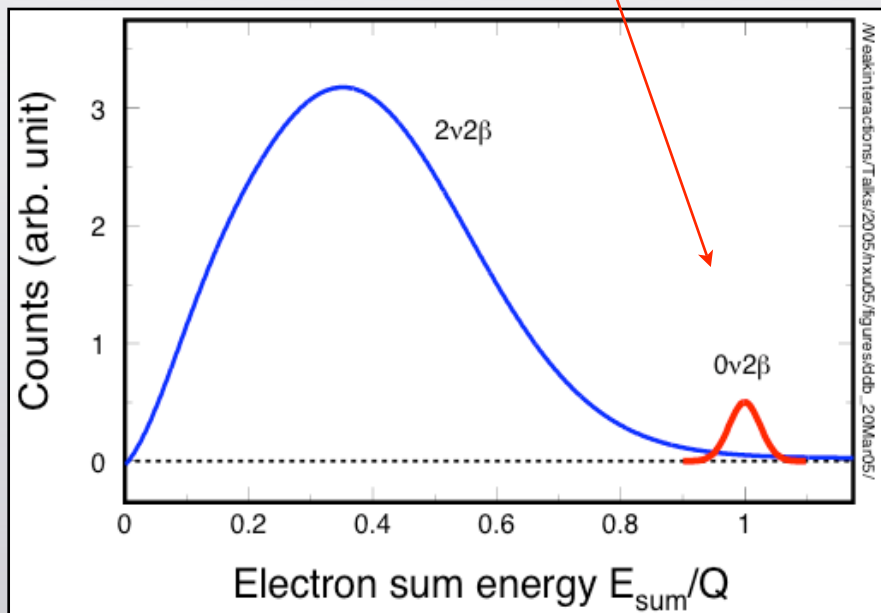




# $\beta\beta 0\nu$ : The measurement



Smearing from energy resolution; introduces background



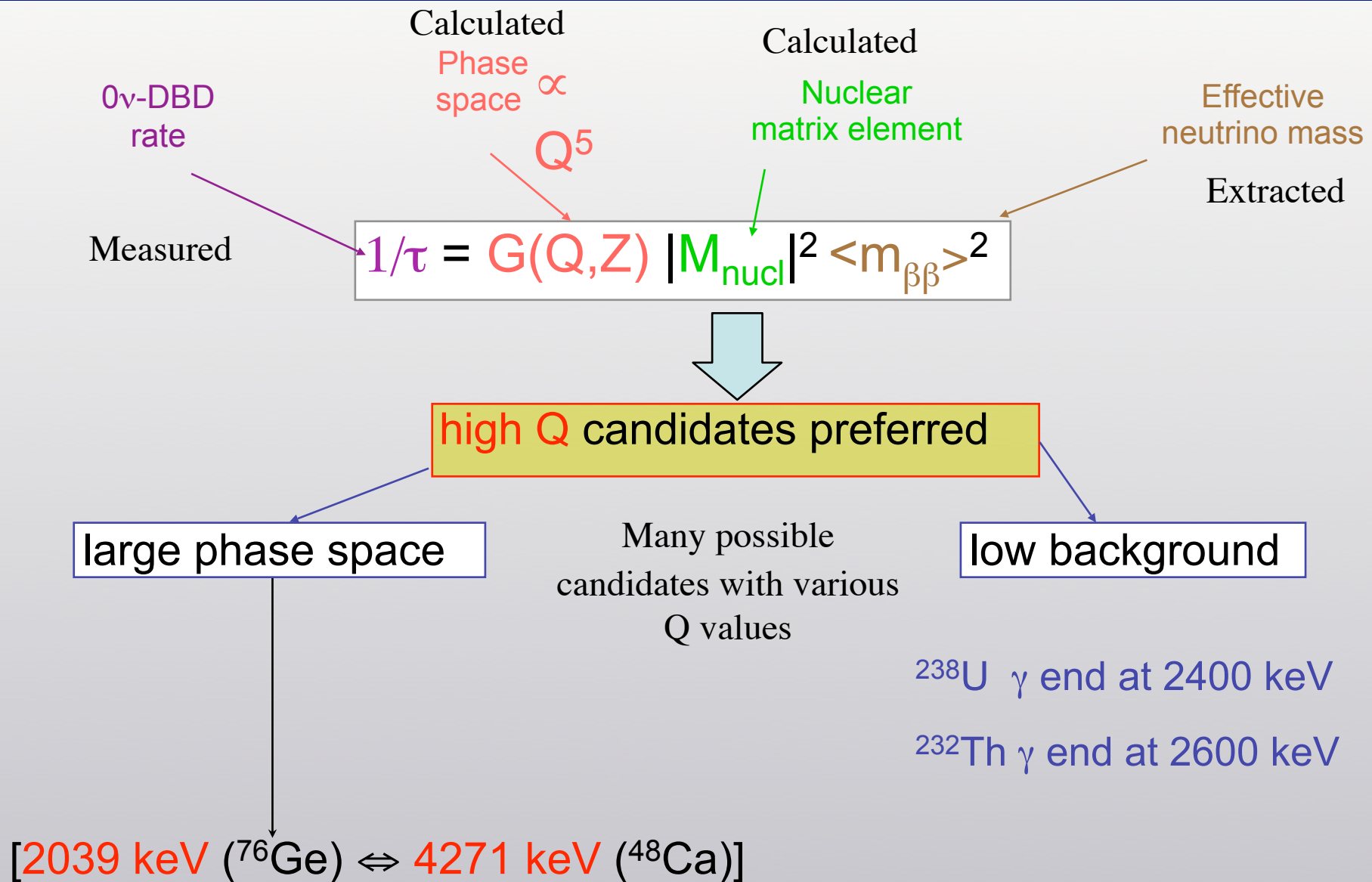
Each candidate isotope has a predicted Q value for the reaction

This is the maximum kinetic energy available to the electron pair

Measure the two electron sum energy:  
The two neutrino mode will be a continuum  
The zero neutrino mode will be a peak

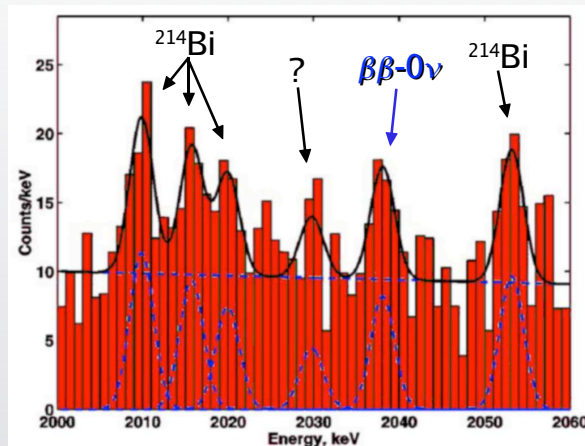


# $\beta\beta 0\nu$ Rate and Neutrino Mass





# $\beta\beta 0\nu$ Scientific Context



The gauntlet is down:  
There is an existing  
 $\beta\beta 0\nu$  claim in  $^{76}\text{Ge}$   
“KKDC”

Klapdor et al., Phys. Lett B 586 (2004) 198

APS Neutrino Study (2004)

We recommend, as a high priority, a phased program of sensitive searches for neutrinoless nuclear double beta decay. In this rare process, one atomic nucleus turns into another by emitting two electrons. Searching for it is very challenging, but the question of whether the neutrino is its own antiparticle can only be addressed via this technique. The answer to this question is of central importance, not only to our understanding of neutrinos, but also to our understanding of the origin of mass.



**Recommendation:** The Neutrino Scientific Assessment Group recommends that the highest priority for the first phase of a neutrino-less double beta decay program is to support research in two or more neutrino-less double beta decay experiments to explore the region of degenerate neutrino masses ( $\langle m_{\beta\beta} \rangle > 100 \text{ meV}$ ). The knowledge gained and the technology developed in the first phase should then be used in a second phase to extend the exploration into the inverted hierarchy region of neutrino masses ( $\langle m_{\beta\beta} \rangle > 10 - 20 \text{ meV}$ ) with a single experiment.

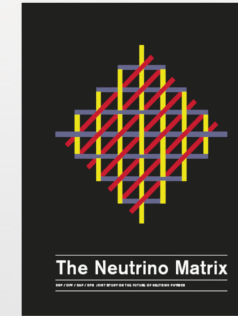
NuSAG report to Nuclear  
Science Advisory  
Committee  
(2005)



# Strategies and Tactics



$$\beta\beta 0\nu$$



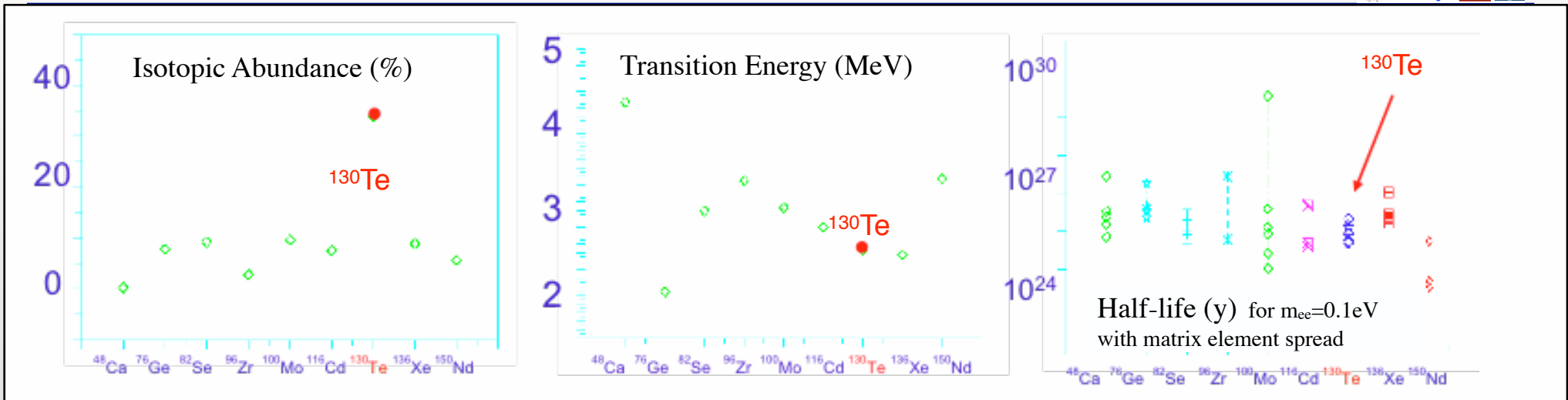
- $\beta\beta 0\nu$  is one of the top priorities in neutrino physics
- Near term:
  - KKDC claim must be addressed
    - Different isotopes
    - Different methods
  - If necessary:
    - Build bigger experiments to gain more sensitivity
      - more mass means more decays in a fixed amount of time

The massive tellurium cryogenic bolometers Cuoricino and CUORE are prepared to play their role this grand neutrino adventure





# Why Tellurium?

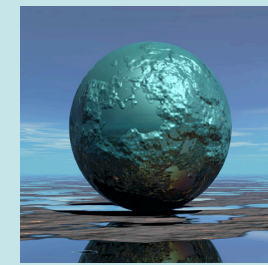


- Cost effective: Enrichment not required
  - Natural abundance 33.87%
- Reasonable  $Q = 2528.8\text{ keV}$ 
  - Large phase space
  - Low gamma background:  $Q$  sits between the Compton edge (2360 keV) and full  $^{208}\text{Tl}$  energy (2615 keV)
- $\beta\beta 2\nu$  observed with geochemical techniques = potential neutrinoless double beta decay candidate

Discovered : by Baron Franz Muller von Reichenstein in 1783  
 Isolated in Sibiu, Romania  
 Origin : The name is derived from the Latin 'tellus', meaning Earth.

“Its compounds are to be avoided because not only are they poisonous but contact with even the tiniest amounts leads to unpleasant body odors!”

$^{130}\text{Te}$



$\text{TeO}_2$  crystals used today in very high end opto-acoustic laser printers for lithography

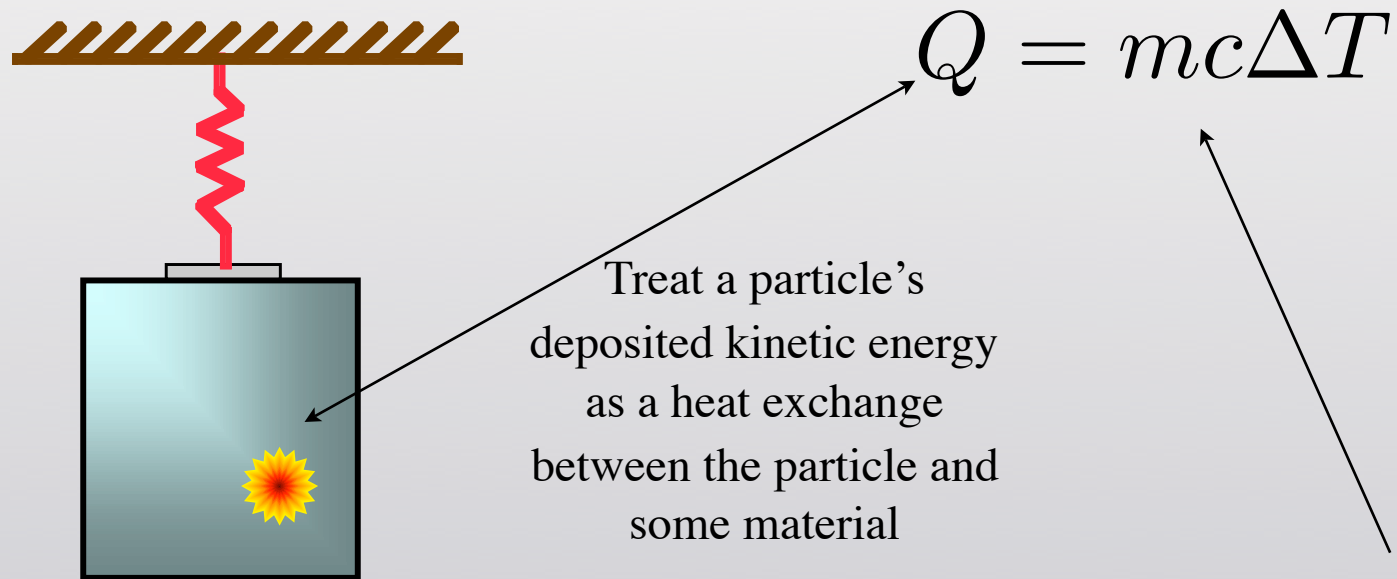
<http://www.chemsoc.org/viselements/>



# Did someone say Bolometer?



- What is a Bolometer?
  - A detector that measures the energy of a particle through calorimetry



Measure the temperature change of the material

To get any measurable temperature rise, the specific heat capacity will have to be insanely small: this means you have to cool things down

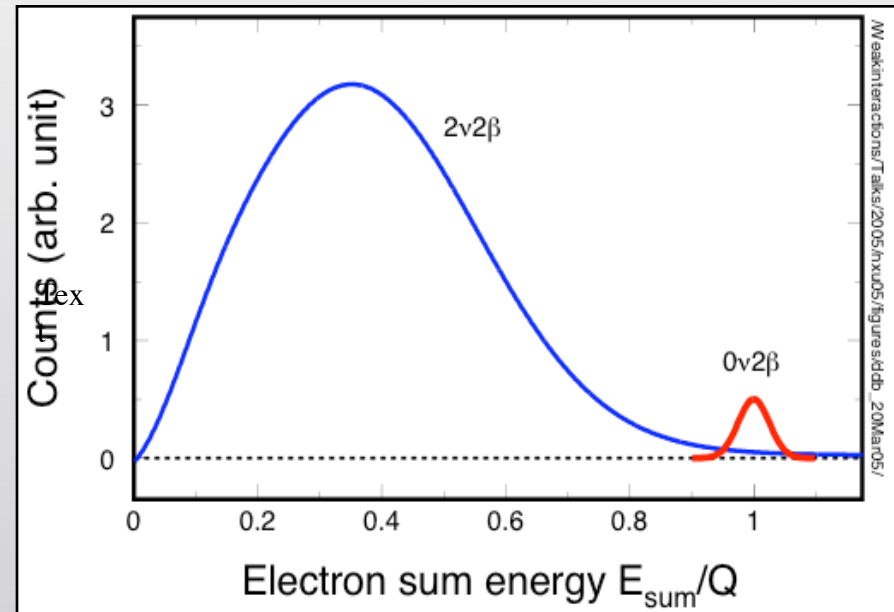
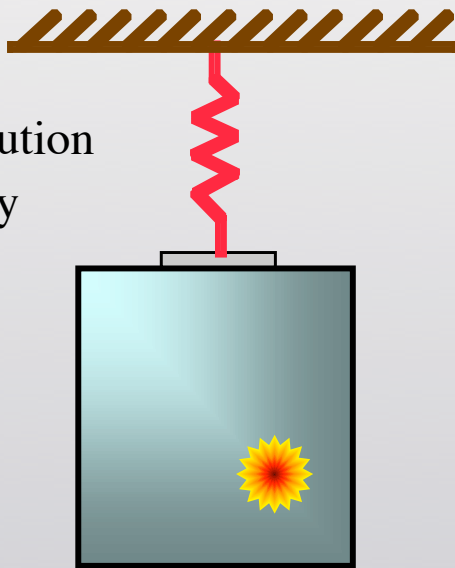


# Source = Detector Bolometry



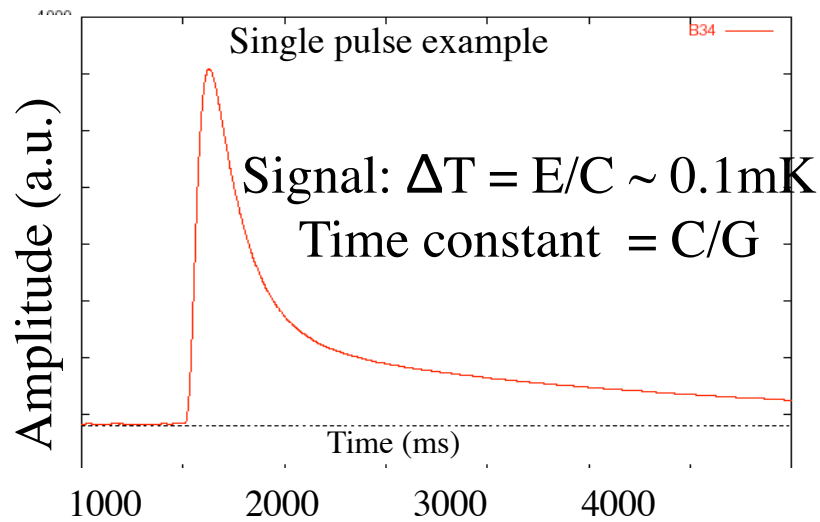
We want to measure the total electron energy of some material undergoing neutrinoless double beta decay: why not build the entire bolometer out of the material that is decaying?

Great energy resolution  
Amazing efficiency



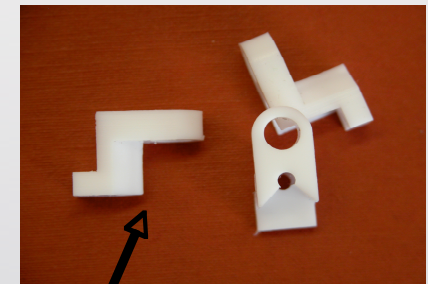
This is exactly what we have done using Tellurium as the source and the detector

# Massive Cryogenic Bolometers



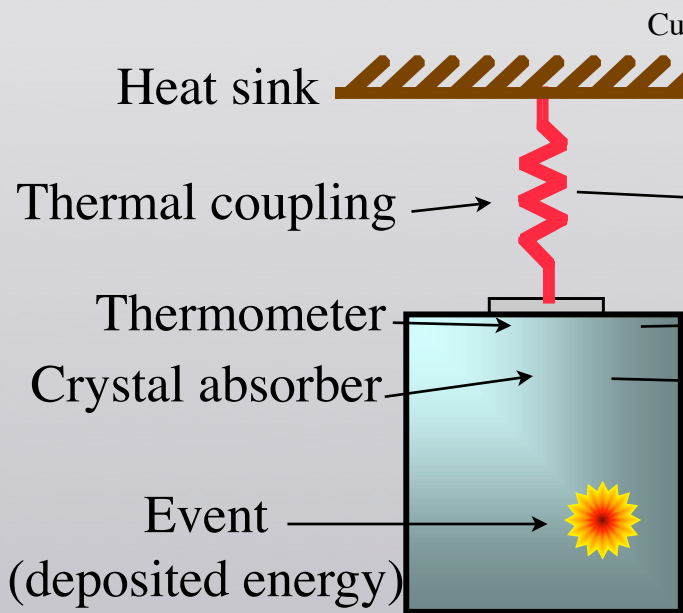
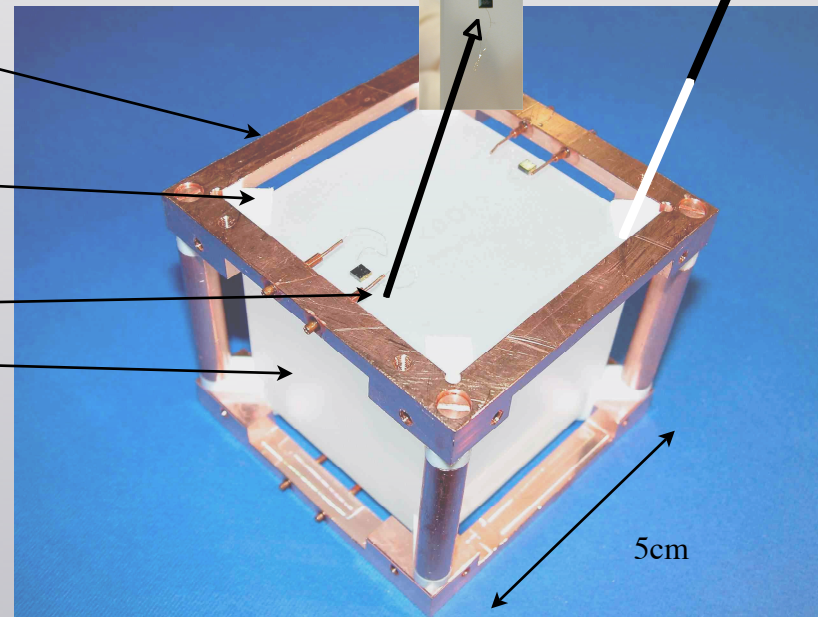
- Dielectric diamagnetic materials
- Low temperatures ( $\sim 10\text{mK}$ )
- Low heat capacity  $\sim T^3$ 
  - $C \sim 2 \text{ nJ/K} = 1 \text{ MeV} / 0.1 \text{ mK}$

**1 mV/1 MeV**



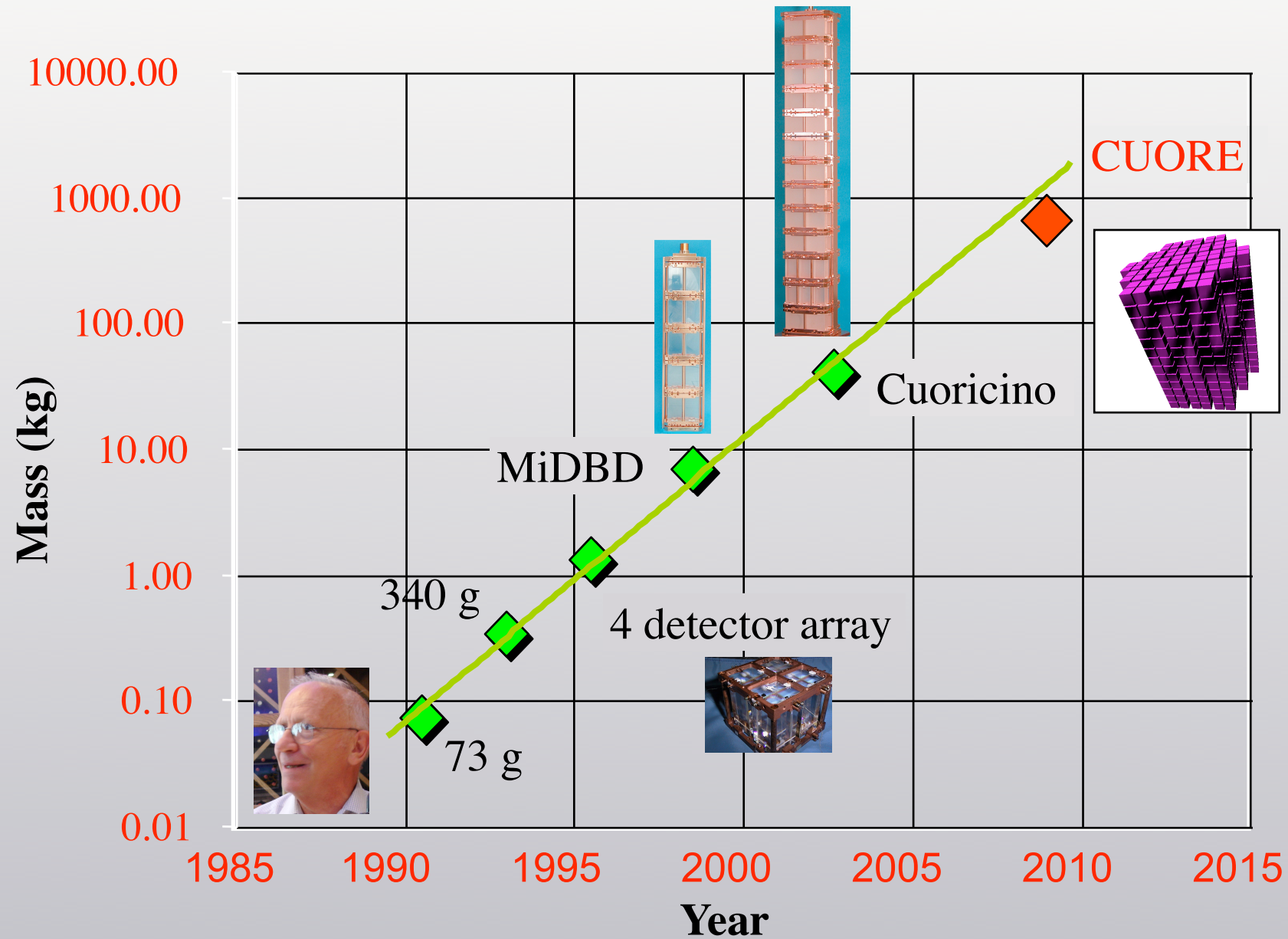
NTD Ge

$G = 4 \text{ pW/mK}$





# “Moore’s Law” Scaling of TeO<sub>2</sub> Bolometry Experiments



# Gran Sasso National Laboratory (LNGS)



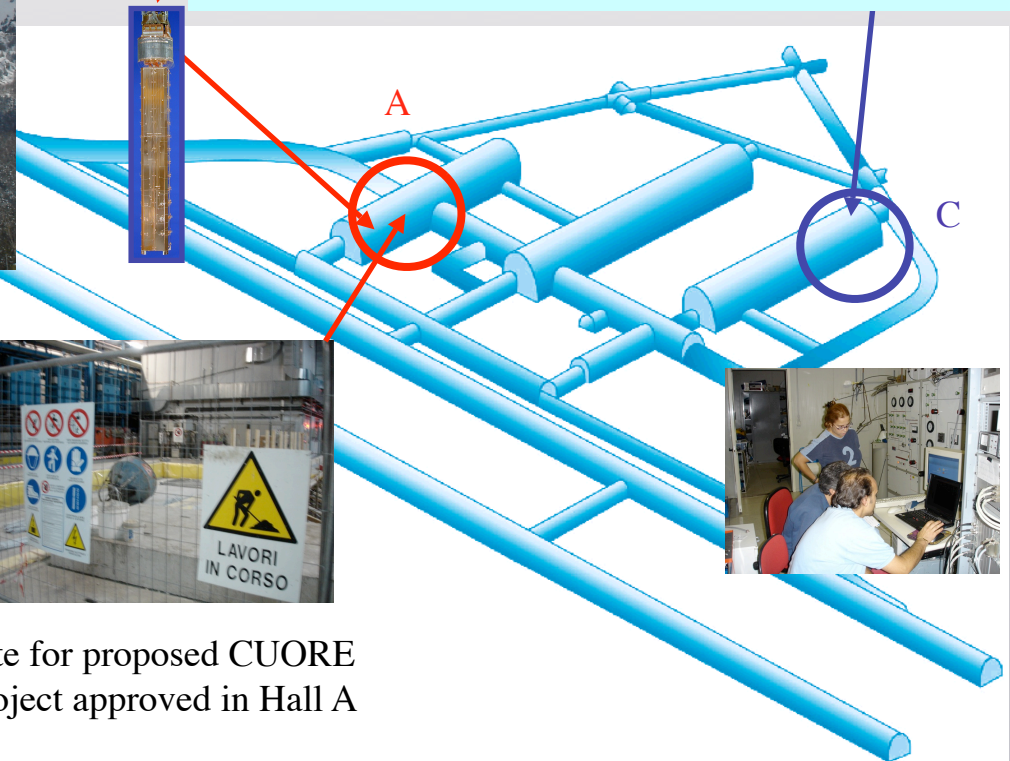
Underground to reduce cosmic rays

Two dilution refrigerators

1. Hall A (Cuoricino) ⇒

Running!

2. Hall C (R&D final tests for CUORE)



Cuoricino is currently the largest operating bolometer and double beta decay experiment in the world

Site for proposed CUORE project approved in Hall A



# Life at the LNGS: External Lab



World Class Research Facility

Beautiful Italian countryside

2 hours East of Rome in the  
Gran Sasso Range







# LNGS: Underground



Many experiments running and/or preparing to run





# Cuoricino, the “little heart” of Gran Sasso

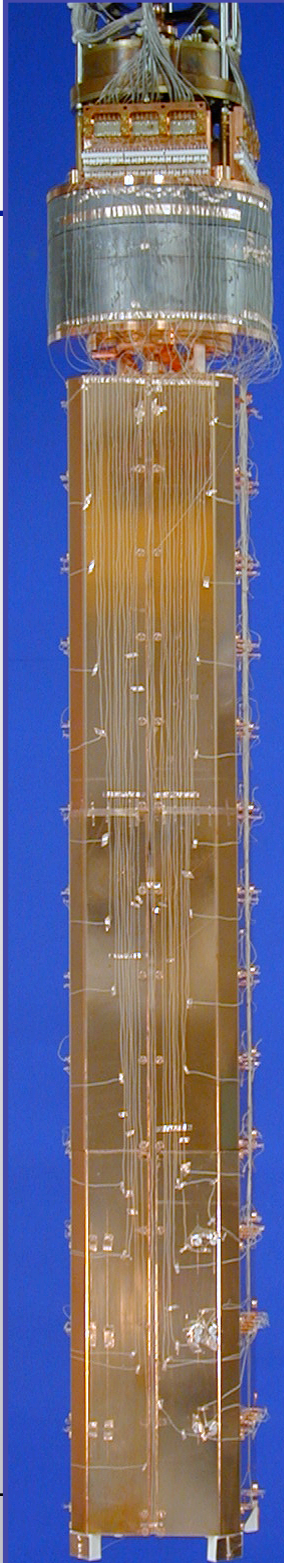
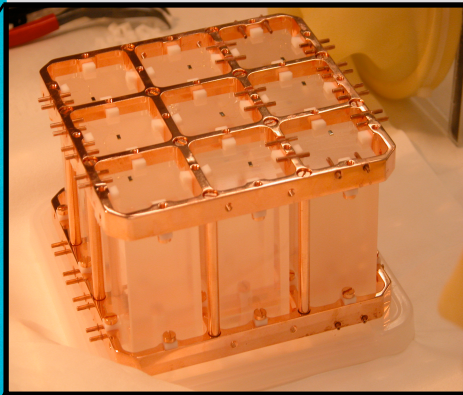
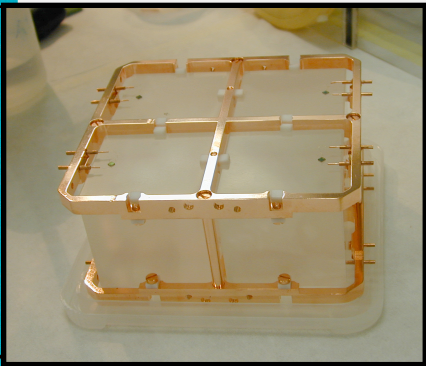
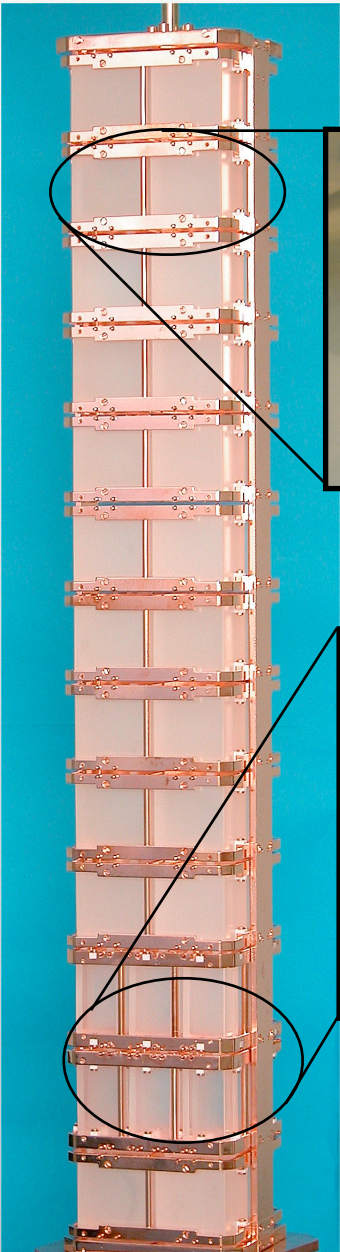
Cooled to 10mK with a powerful dilution refrigerator

11 modules, 4 detector each,  
crystal dimension:  $5 \times 5 \times 5 \text{ cm}^3$   
crystal mass: 790 g  
 $44 \times 0.79 = 34.76 \text{ kg of TeO}_2$

Encased in a cryostat, lead shield, nitrogen box, neutron shield, and Faraday cage

2 modules x 9 crystals each  
crystal dimension:  $3 \times 3 \times 6 \text{ cm}^3$   
crystal mass: 330 g  
 $18 \times 0.33 = 5.94 \text{ kg of TeO}_2$

Total detector mass:  $40.7 \text{ kg TeO}_2 \Rightarrow 11.34 \text{ kg } ^{130}\text{Te}$





# Cuoricino Results from Runs 1&2: No Peak

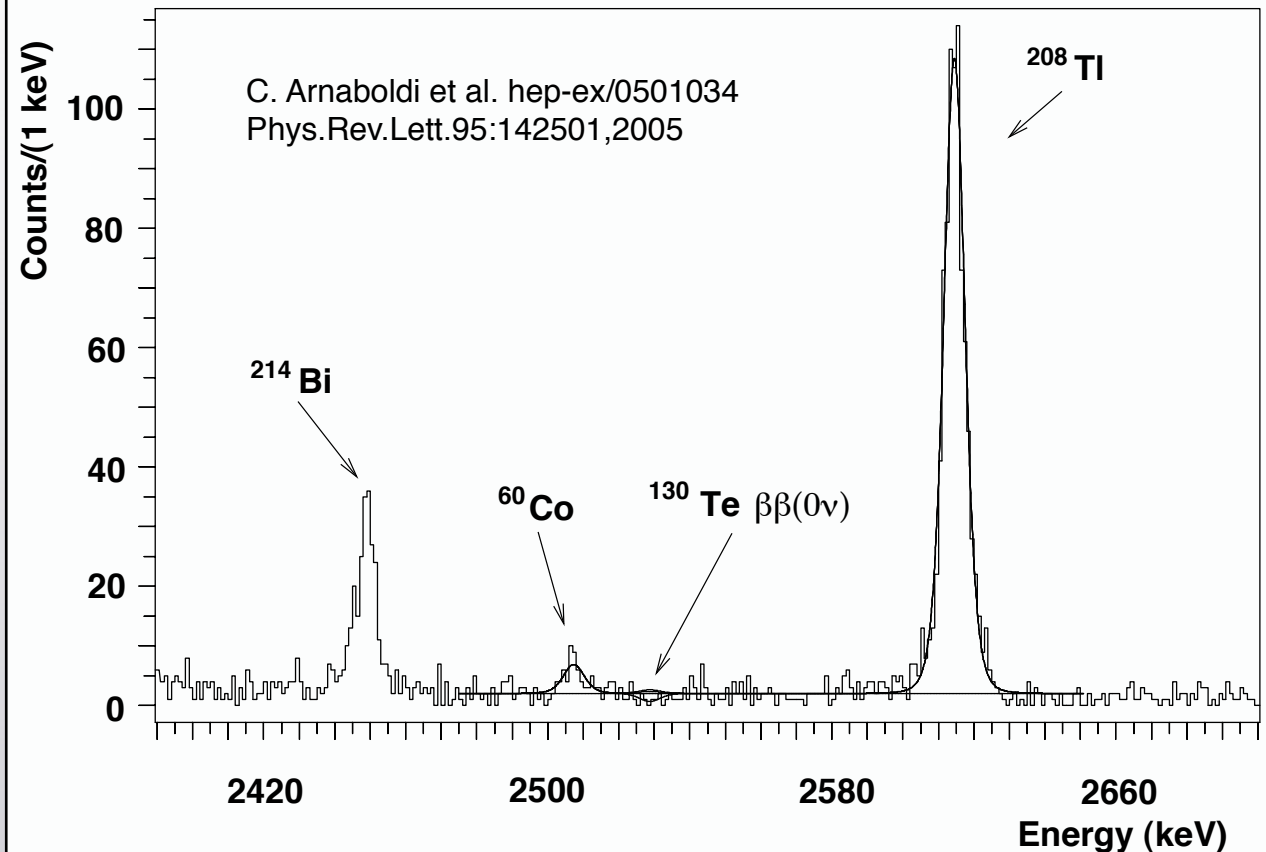


Exposure  
= 10.85 kg y

Resolution:  
FWHM at 2615 keV  
=  $9.2 \pm 0.5$  keV

Background:  
In the  $\beta\beta 0\nu$  region  
=  $0.18 \pm 0.01$  counts / (keV kg y)

- Anticoincidence spectrum
- $\epsilon \sim 85\%$
- Maximum Likelihood + Flat +  $^{60}\text{Co}$  (2505 keV)
  - energy region: 2480-2650 keV
- $\sim 5\%$  systematic error upon parameter variation
- N-Gaussian response function with individual FWHM detector resolution @ 2615 keV



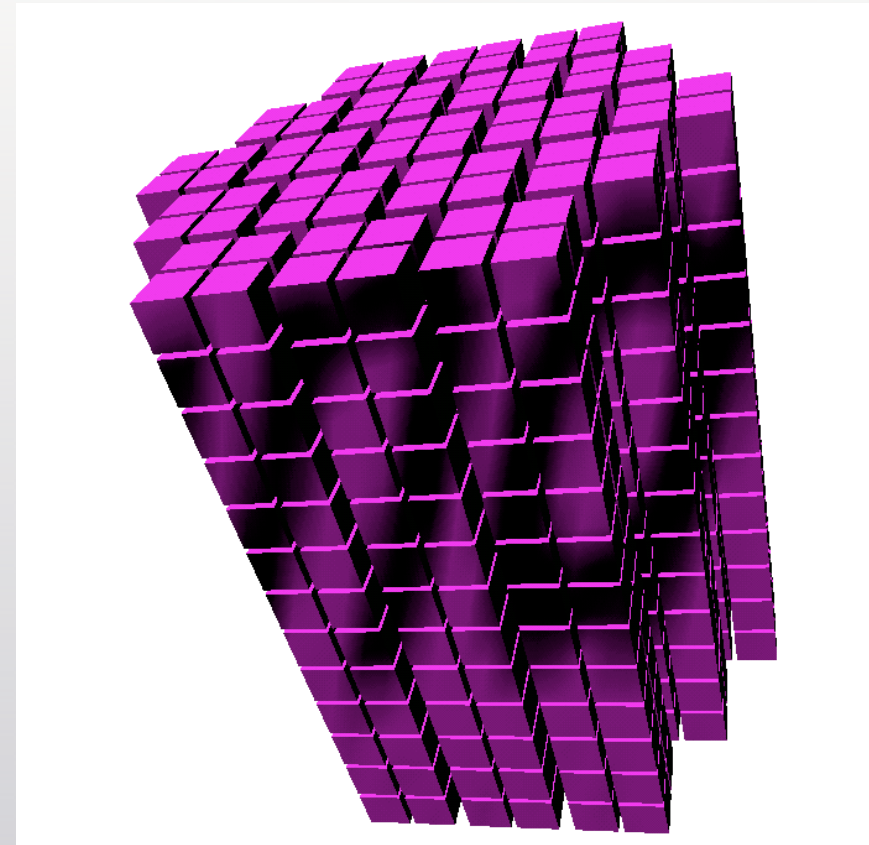
No peak found  
 $\tau^{0\nu}_{1/2} > 1.8 \times 10^{24}$  y at 90% C.L.  
 $m_\nu < 0.2 - 1.1$  eV

Spread is due to a range of published matrix elements

# CUORE: Cryogenic Underground Observatory for Rare Events



- Array of 988  $\text{TeO}_2$  crystals
- 19 Cuoricino-like towers suspended in a cylindrical structure
- 13 levels of 4  $5 \times 5 \times 5 \text{ cm}^3$  crystals (750g each)
- $^{130}\text{Te}$ : 33.8% isotope abundance
- Time of construction: 4 years
- Total cost: 14-17M USD (depends on Euro...)
- 1st Data target: Jan 1, 2010



$750 \text{ kg TeO}_2 \Rightarrow 200 \text{ kg } ^{130}\text{Te}$

Acts as a single,  
highly segmented,  
detector

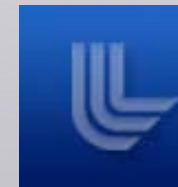
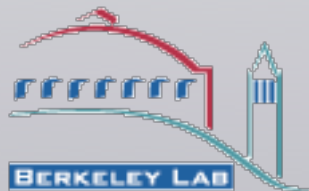
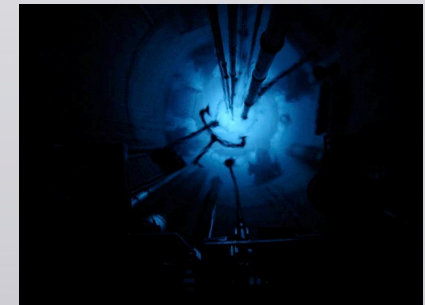
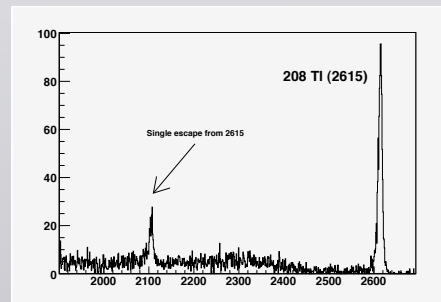
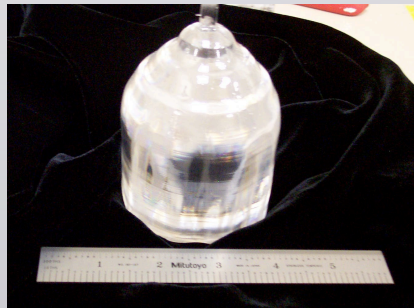
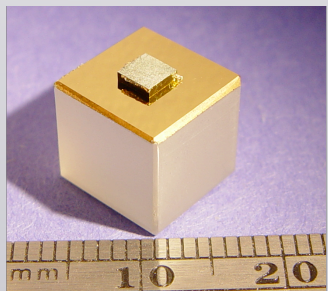
Approved by the Science Counsel of Gran Sasso Laboratory and by INFN



# LBL and LLNL Efforts in CUORE and Cuoricino



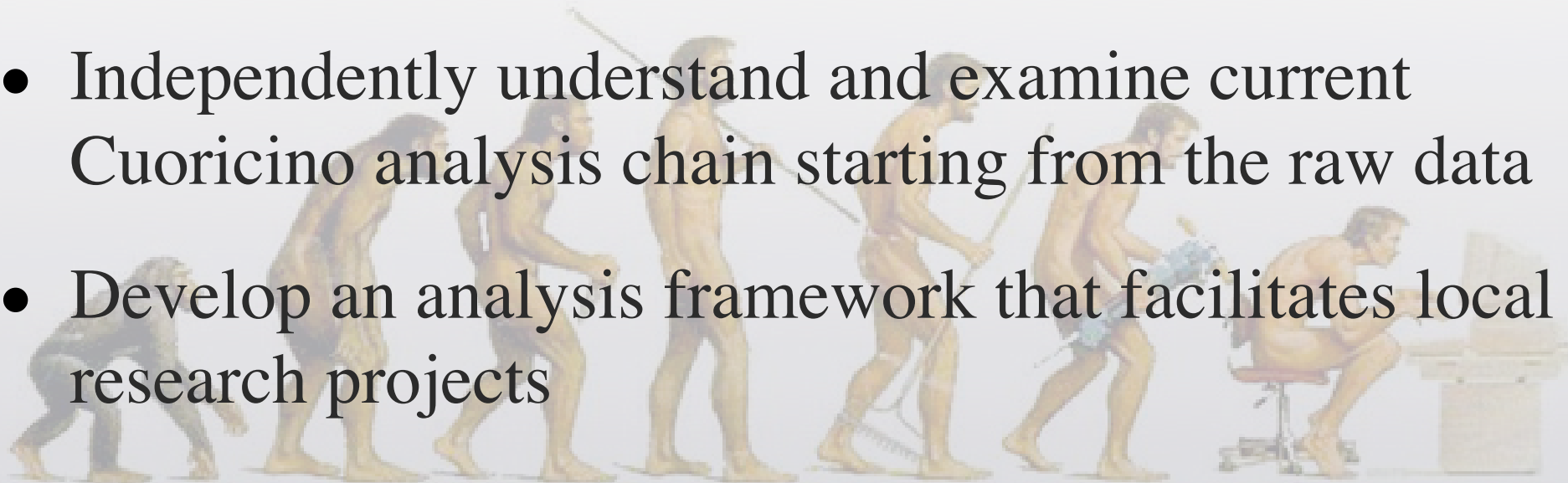
- Thermistor Production
- Crystals
- Data Analysis
- Background Studies
- Construction



# My Analysis Efforts at LBL



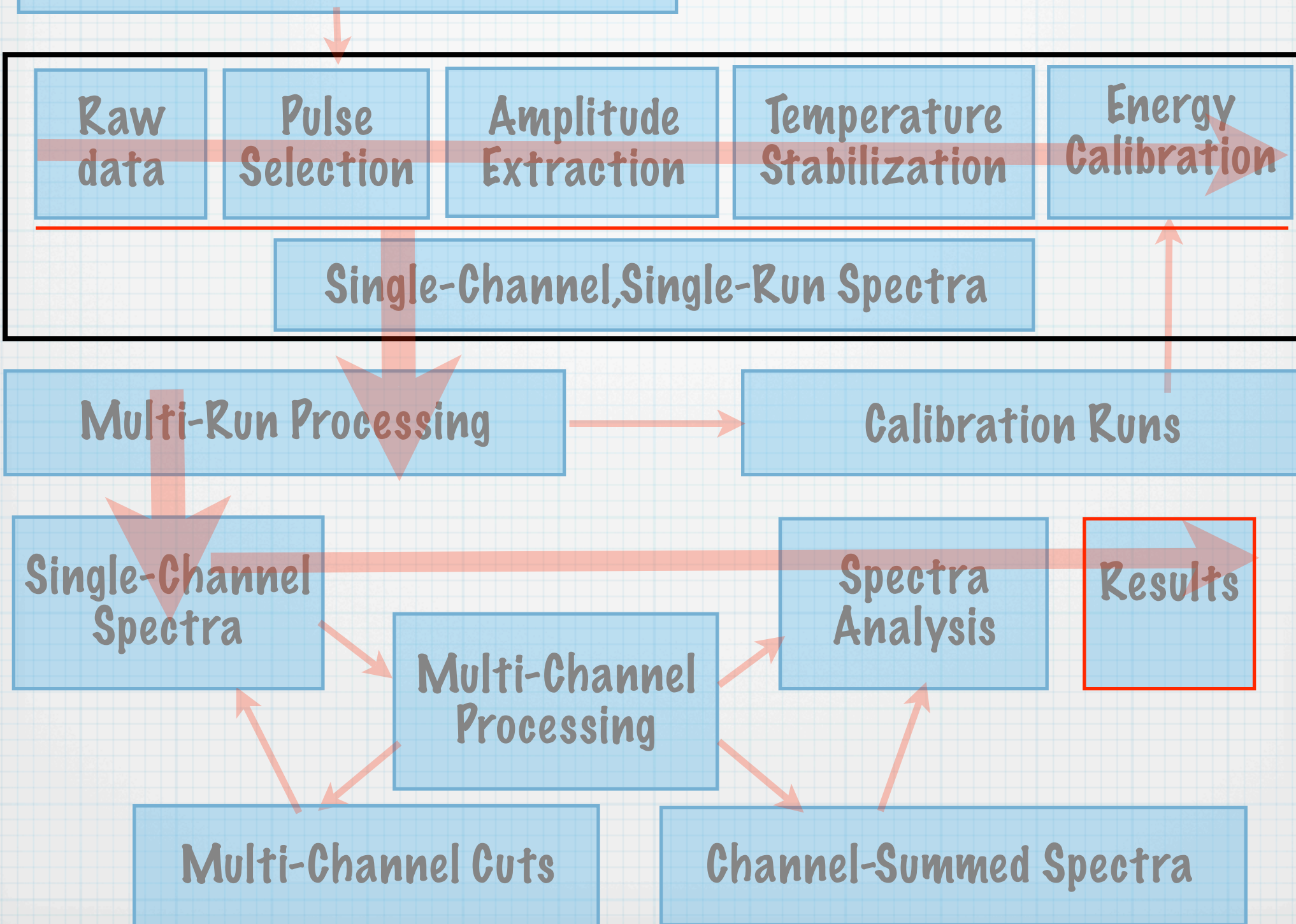
- Independently understand and examine current Cuoricino analysis chain starting from the raw data
- Develop an analysis framework that facilitates local research projects
- Contribute collaboratively to ongoing CUORE data analysis development and standards



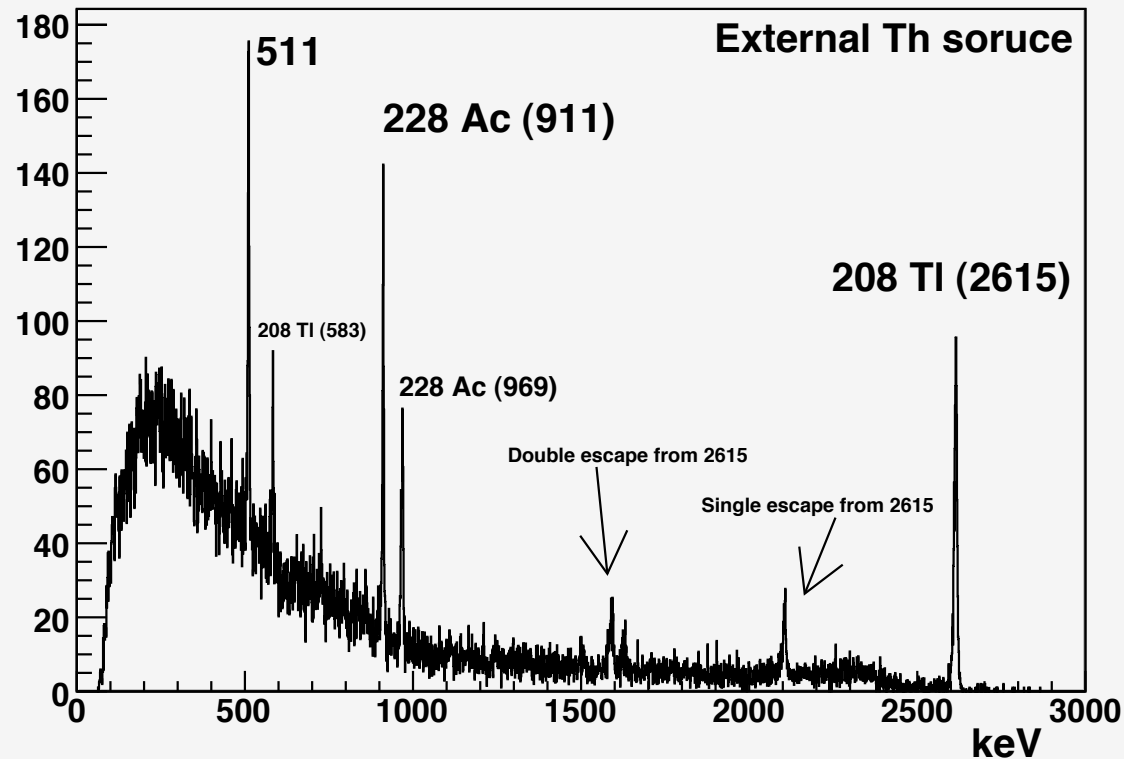


# Run and Channel Pre-selection

# Schematic Analysis Chain



# Berkeley Analysis Output



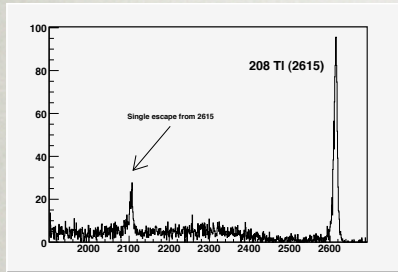
- Calibration Data
- Independent “first principles” spectra from raw data

- Work with the raw data
- Understand nuclear process and backgrounds
- Understand the detector
- Lots of fun programming (C++ etc.)

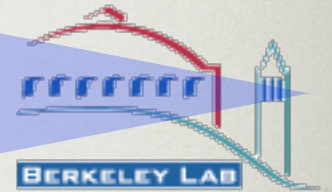
Many possible  
fun senior projects



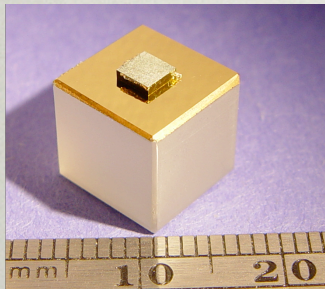
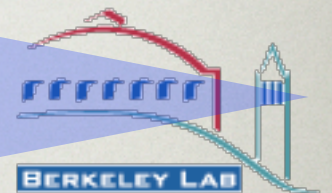
# POTENTIAL LOCAL STUDENT PROJECTS



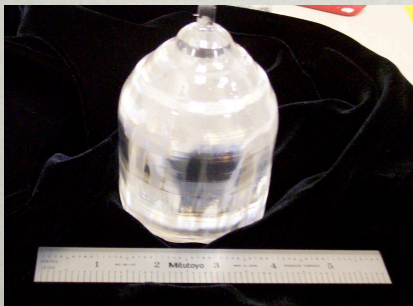
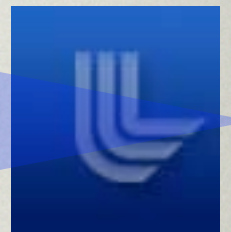
Cuoricino and CUORE  
Data Analysis (my focus)  
Perhaps eventually set up a small  
dilution refrigerator?



Background analysis



Thermistor and Misc R&D



Crystals





# Shifts at Gran Sasso for Cuoricino



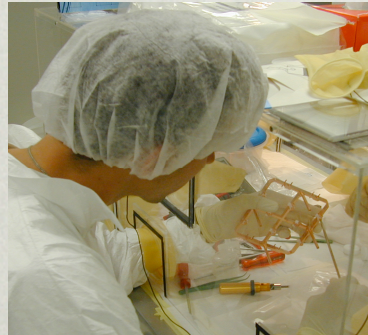
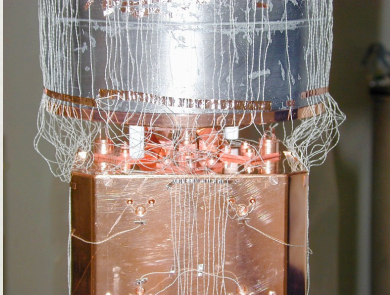
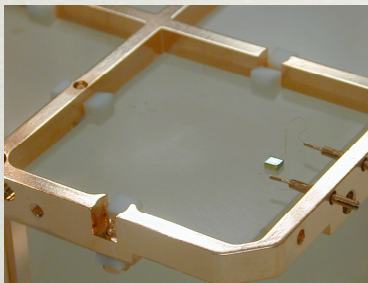
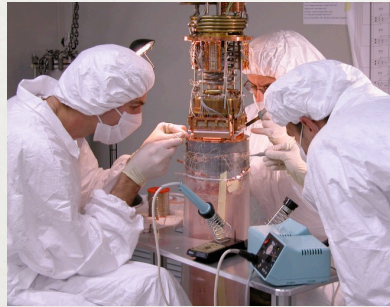
Spend several weeks  
in Gran Sasso  
during the summer  
working on  
Cuoricino

The work will be “on  
the job training” for  
physics students at  
all levels



# REMOTE STUDENT PROJECTS

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In the coming years there will be plenty of exciting work at Gran Sasso to be done in preparing for CUORE

The work will be hands-on “on the job training” for physics students at all levels

Many potential fun hands-on senior projects



# SUMMARY

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- Neutrinoless Double Beta Decay can help determine neutrino's mass and nature
- CUORE and Cuoricino are cutting edge calorimetry experiments poised to measure this process
- Plenty of room for student involvement on many levels
- Exciting time for neutrino physics!



# The CUORE Collaboration



## CUORE author list

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# FURTHER READING AND VIEWING

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- The Neutrino Matrix [physics/0411216](#)
- CUORE proposal [hep-ex/0501010](#)
- CUORE website <http://crio.mi.infn.it/wig/>, <http://cuore.lbl.gov/>
- The Ghost Particle (NOVA) <http://www.pbs.org/wgbh/nova/neutrino/>
- Particle data group <http://pdg.lbl.gov/>



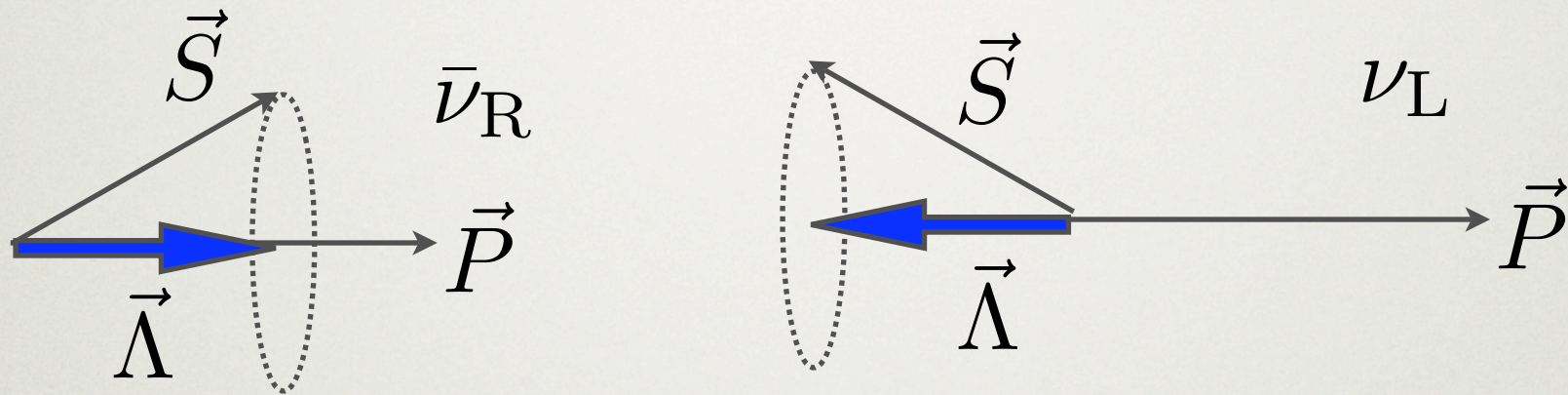
# BACKUP

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# THE ANTINEUTRINO

Can we really tell neutrinos apart from their antiparticles?



The weak force, when interacting with neutrinos, only seems to interact with particles we call “left-handed neutrinos” or “right-handed antineutrinos”

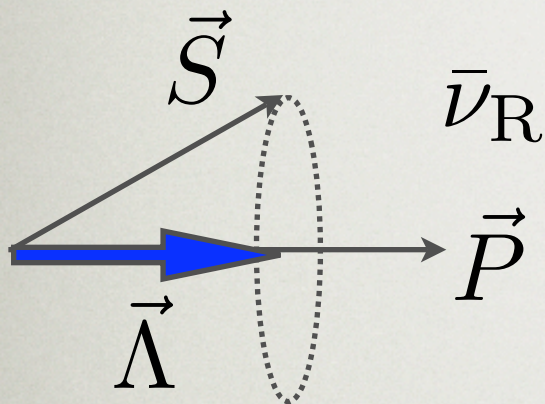
“Handedness” is then regarded as the projection of the spin along the momentum vector

If the neutrino is massless it is moving at the speed of light; this categorization works well because the handedness is fixed; we simply accept that “other” particles don’t exist for some reason

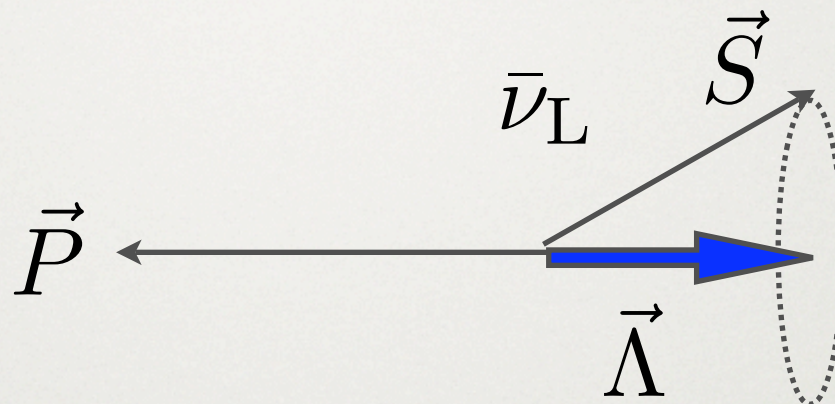


# THE ANTINEUTRINO

Can we really tell neutrinos apart from their antiparticles?



In our frame:  
moving slower than the  
antineutrino



Same particle in the frame of the little purple running dude:  
moving faster than the antineutrino



If the neutrino has mass we should be able to convert it into particles  
we thought didn't exist just by changing reference frames!

What if the objects we called the "neutrino" and "antineutrino" were  
just different spin projections of the same kind of particle all along?

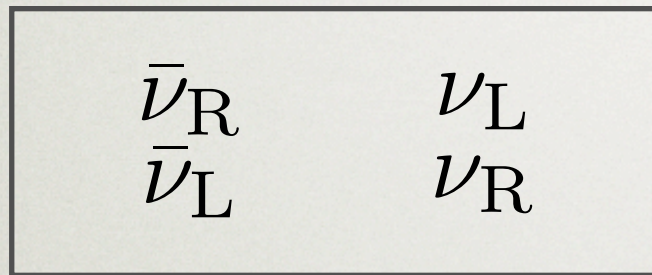


# THE ANTINEUTRINO

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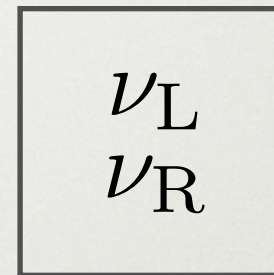
Can we really tell neutrinos apart from their antiparticles?

Dirac



Four particles only two are observed

Majorana

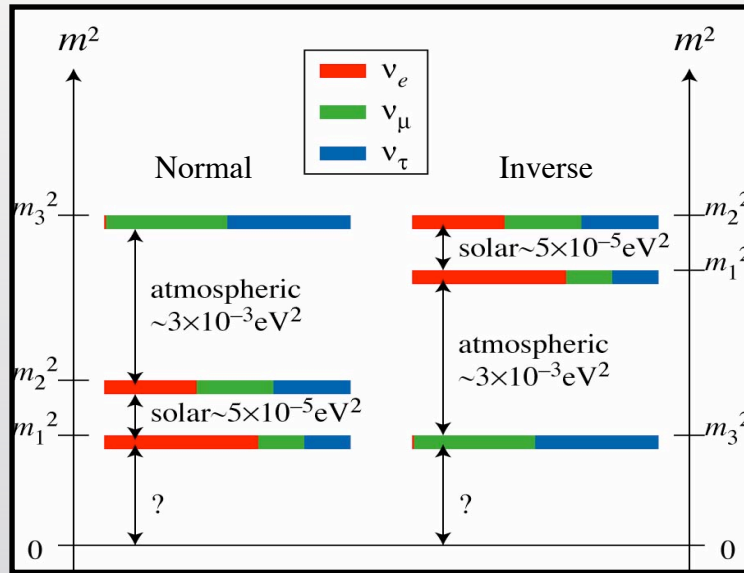


Two particles, both are observed;  
regarded as “more natural” for  
massive neutrinos

I will now refer to this issue as a question of the neutrino’s “nature”

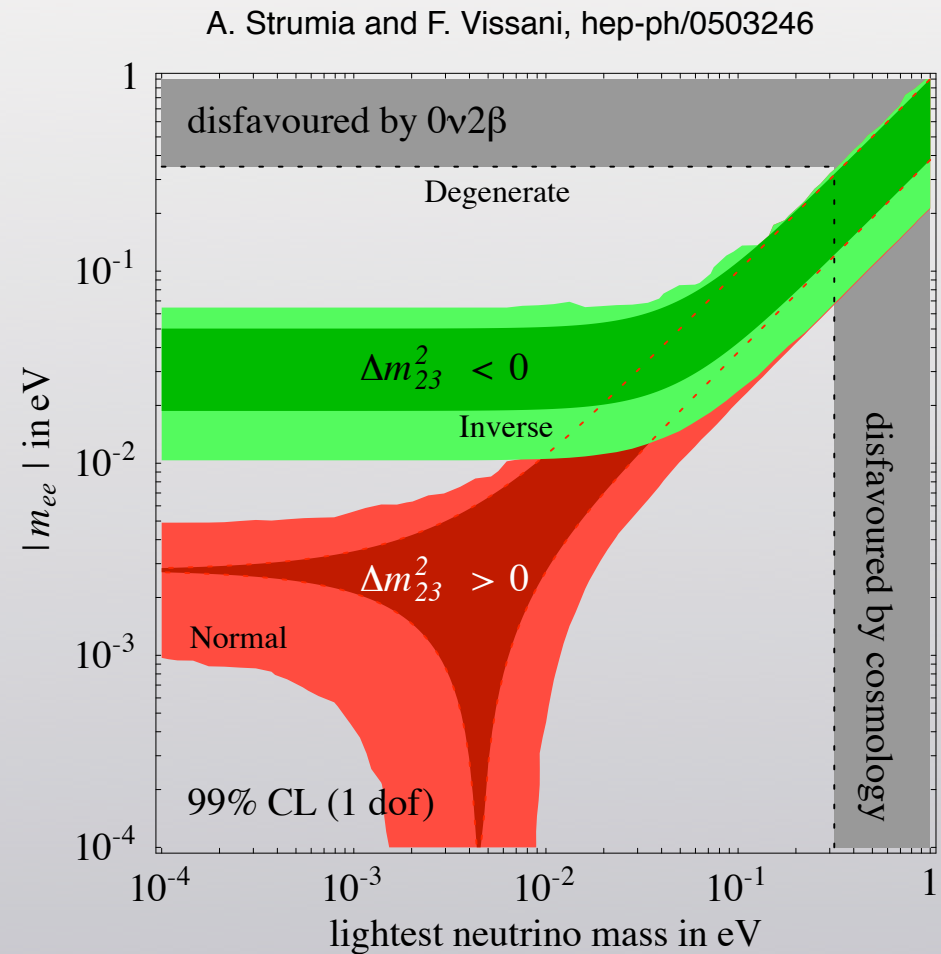


# $\beta\beta 0\nu$ : a powerful observational tool



$$1/\tau = G(Q,Z) |M_{\text{nucl}}|^2 \langle m_{\beta\beta} \rangle^2$$

The rate depends on the number of moles of active material available



# CUORE Sensitivity



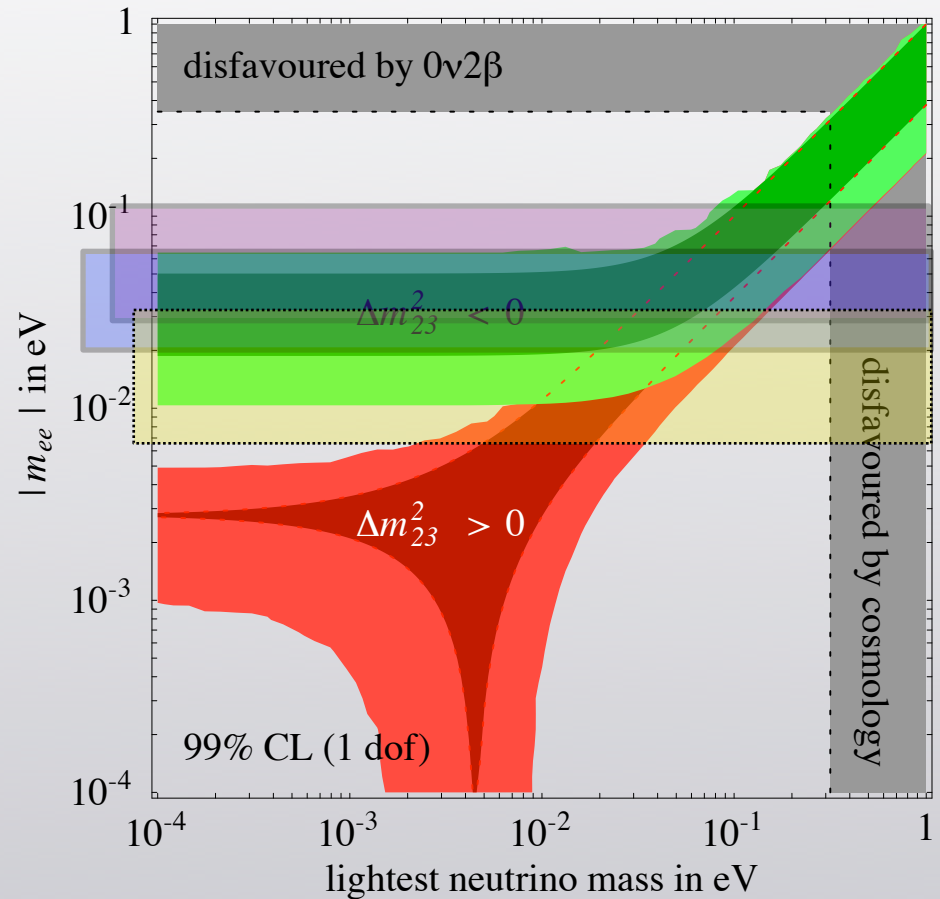
Five year sensitivity based on detector resolution, background, and matrix element spread

B(counts/keV/kg/y)	$\Delta$ (keV)	$T_{1/2}$ (y)	$ \langle m_{\nu} \rangle $ (meV)
0.01	10	$1.5 \times 10^{26}$	23–118
0.01	5	$2.1 \times 10^{26}$	19–100
0.001	10	$4.6 \times 10^{26}$	13–67
0.001	5	$6.5 \times 10^{26}$	11–57

More optimistic but plausible case: eliminate degenerate hierarchy and continue excavation deeper into inverse hierarchy

Fantasy: 99% enriched CUORE after 10 years with 5 keV resolution and 0.001 c/keV/kg/y

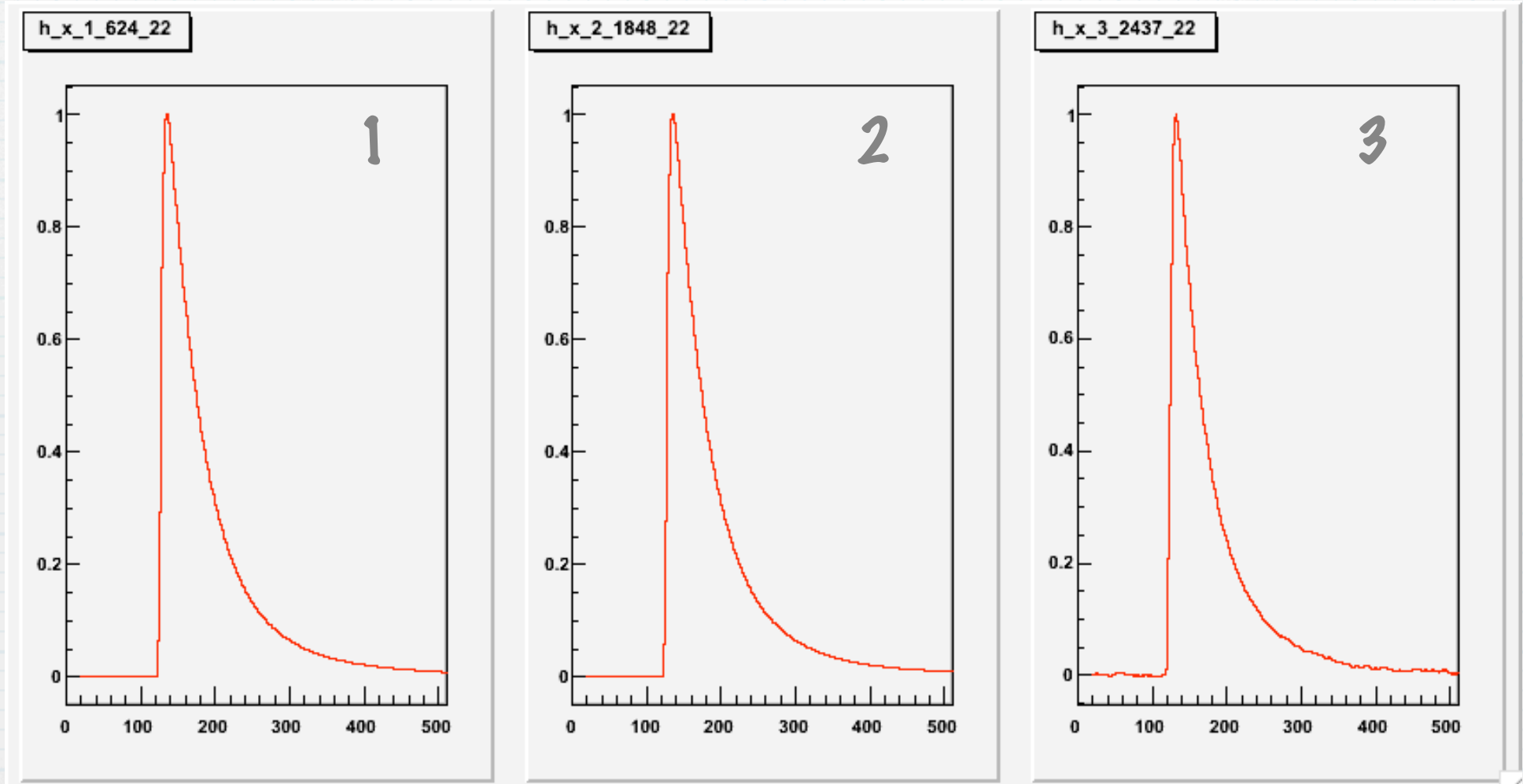
A. Strumia and F. Vissani, hep-ph/0503246



10 years from now, with input from different isotopes, the theoretical mass spread will hopefully be smaller



# Pulse Selection and pattern recognition software: Simple Example

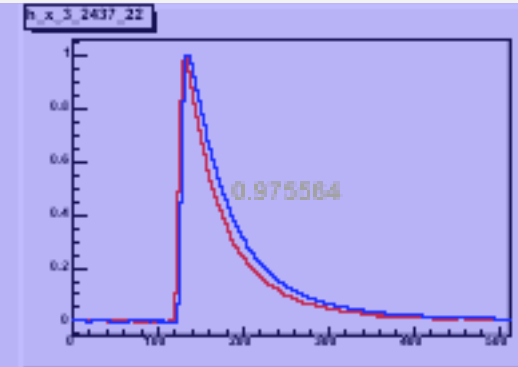
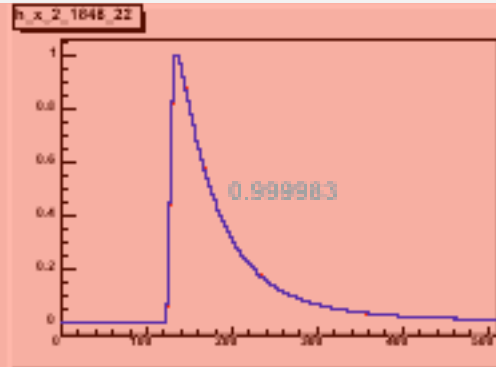
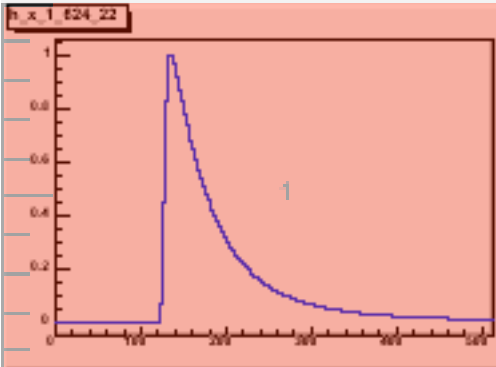


3 pulses scaled to unit height to highlight shape

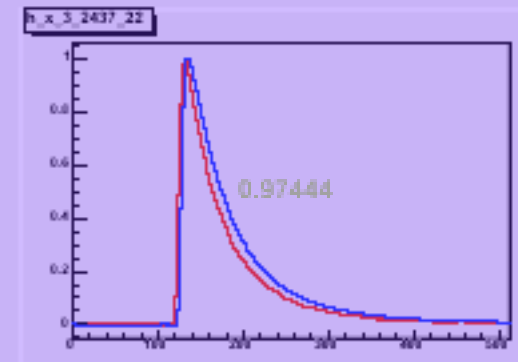
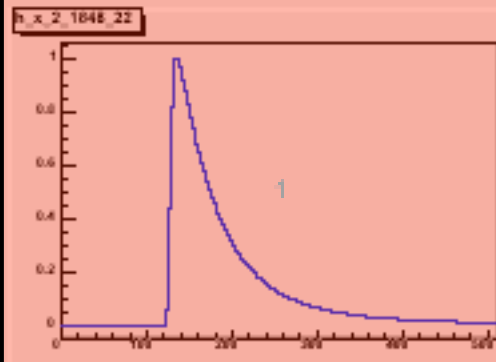
$$C_{ij} = \hat{P}_i \cdot \hat{P}_j$$

[(1,1)=1.000000] [(1,2)=0.999983] [(1,3)=0.975564]  
[(2,2)=1.000000] [(2,3)=0.974440]  
[(3,3)=1.000000]

1



2



3

3 is different than 1 and 2

1

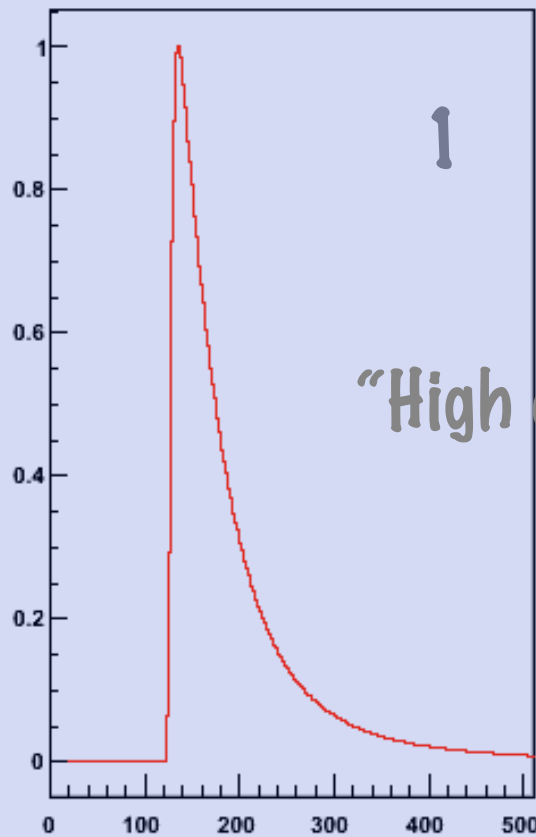
2

3



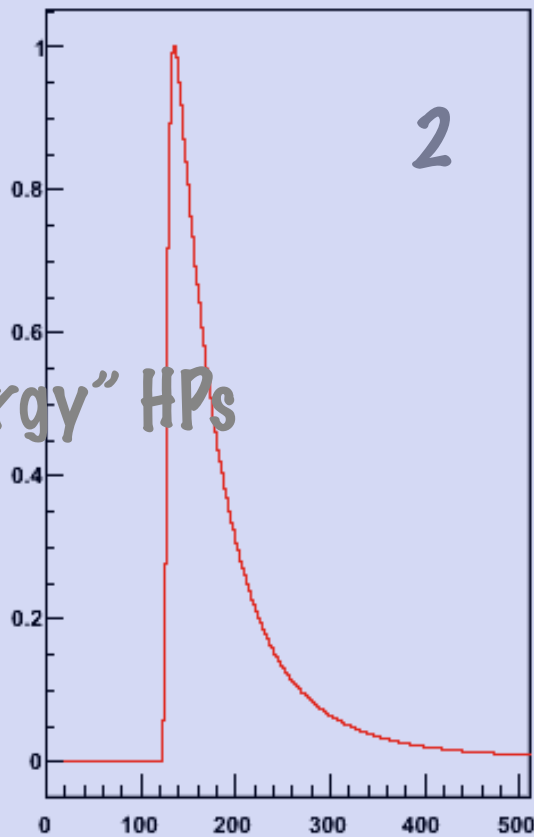


h\_x\_1\_624\_22

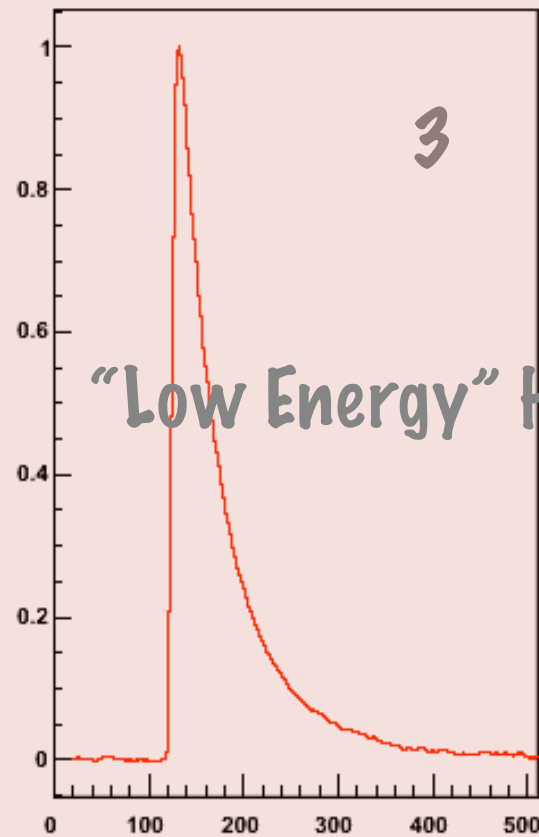


“High energy” HPs

h\_x\_2\_1848\_22



h\_x\_3\_2437\_22



“Low Energy” HP

\*\*\*\*\*

Entry	1
Entry in TTree (1st=0)	624
Channel (tower)	22
Pulse Time (s)	166.434
Rise time t-to-m (s)	0.112
Decay const (~1/s)	0.615127
Max Loc (ch,s)	135.5 , 1.084

\*\*\*\*\*

Entry	2
Entry in TTree (1st=0)	1848
Channel (tower)	22
Pulse Time (s)	471.954
Rise time t-to-m (s)	0.112
Decay const (~1/s)	0.615184
Max Loc (ch,s)	135.5 , 1.084

\*\*\*\*\*

Entry	3
Entry in TTree (1st=0)	2437
Channel (tower)	22
Pulse Time (s)	624.53
Rise time t-to-m (s)	0.096
Decay const (~1/s)	2.33407
Max Loc (ch,s)	131.5 , 1.052